

# Cosmology Assignment 3

## Weeks 5 & 6: Cosmic Structure and Cosmological Observations

Q1) In a Universe with  $\Omega_m = 0.3$  and  $H_0 = 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$ , an early galaxy at  $z = 20$  can be considered as a sphere containing total mass  $10^{10} M_\odot$ , which is a factor 200 more dense than the Universe as a whole.

- What is the radius of this proto-galaxy?
- Estimate the dynamical formation time of the galaxy.
- Assuming the Universe is matter-dominated, by what redshift will the galaxy have formed?

Q2) The galaxy M81 at 3.5 Mpc is one of the nearest galaxies not bound to the “Local Group”. Assuming a cosmology with  $\Omega_m = 0$  and  $\Omega_\Lambda = 1$ , which is a good approximation of our own Universe in the distant future, show that in 100 billion years’ time, light from M81 will not be able to reach us! You can assume  $H_0 = 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$ .

Q3) According to a slightly weird theory, cold dark matter consists of lead balls the size of the Earth randomly distributed across the Universe ( $\rho_{\text{Lead}} = 11,300 \text{ kg m}^{-3}$  and  $R_{\text{Earth}} = 6378 \text{ km}$ ).

- If  $\Omega_{DM} = 0.25$ , how many balls will there be per cubic Mpc?
- If the lead balls are optically thick, what would be the mean free path of light through the Universe? Can we detect the dark matter in this way?
- How would the mean free path depend on the radius of the balls, and what radius would make the Universe opaque to light?
- Why is this not a good model for dark matter?

Q4) The process where electrons and positrons annihilate to form photons ( $e^+ + e^- \leftrightarrow \gamma + \gamma$ ) “freezes out” when the temperature of the Universe is  $T = 5 \times 10^9 \text{ K}$ . Assuming that the Universe is radiation-dominated at this time, calculate the age and redshift of this epoch. You can assume the radiation temperature today is 2.73 K and corresponds to a density  $\Omega_r = 5.1 \times 10^{-5}$ .