

# Class 6: Relativistic Phenomena

*In this class we will apply the laws of relativistic mechanics in a range of settings, and also consider the relativistic Doppler effect*

# Class 6: Relativistic Phenomena

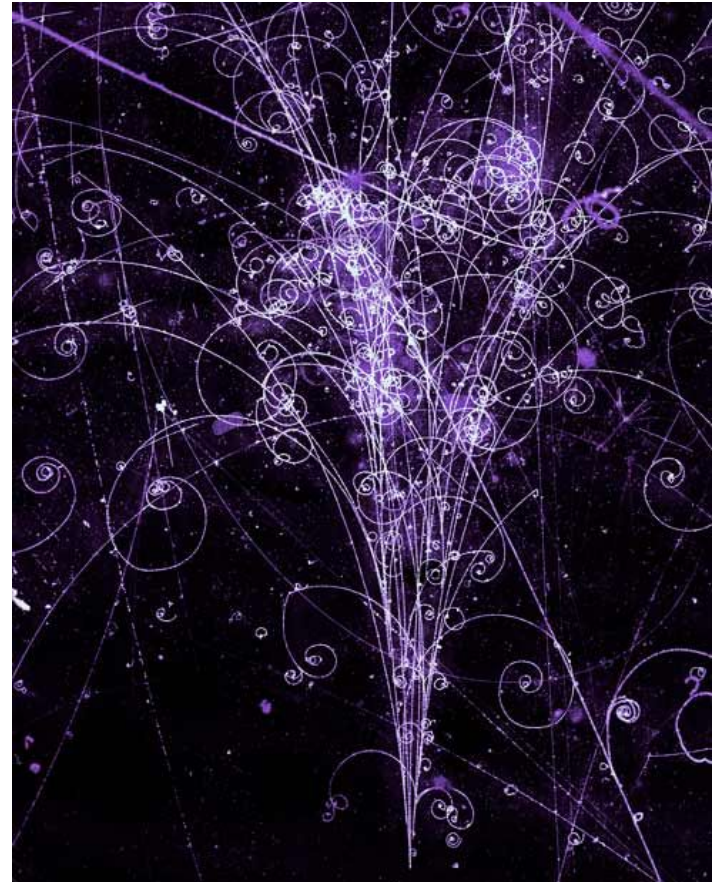
At the end of this session you should be able to ...

- ... recall the relations that apply in special relativity between **energy, momentum and rest mass**, of particles and photons
- ... apply mass-energy equivalence to determine the energy released in **nuclear fission and fusion**
- ... apply the conservation of relativistic energy and momentum in **interactions, collisions and annihilations** involving matter and radiation
- ... be familiar with the **relativistic Doppler effect**, and the differences compared with the effect at low velocities

# Summary of relativistic mechanics

- *Newtonian mechanics is modified at relativistic speeds*

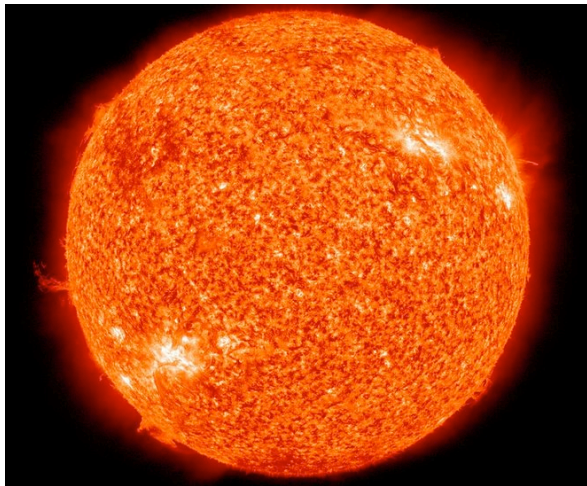
- Relativistic **momentum**,  $p = \gamma_u m_0 u$ , is conserved  $\left[ \gamma_u = 1/\sqrt{1 - \frac{u^2}{c^2}} \right]$
- Relativistic **energy**,  $E = \gamma_u m_0 c^2$ , is conserved
- $E^2 = (pc)^2 + (m_0 c^2)^2$
- There is an equivalent **rest-mass energy** when  $u = 0$ ,  $E = m_0 c^2$
- **Photons** have zero rest mass,  $E = pc$ , where  $E = hf = hc/\lambda$



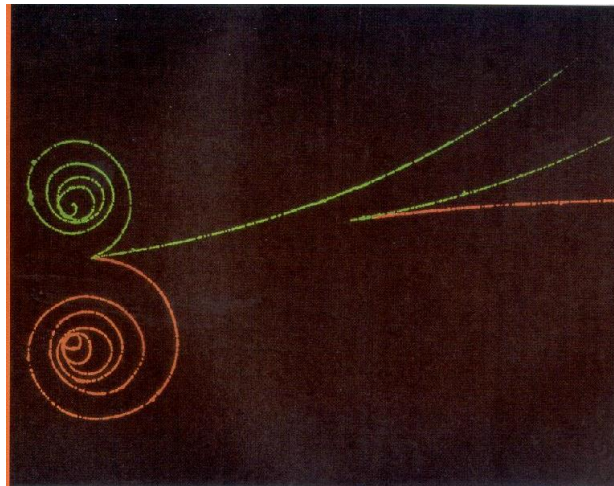
# Relativistic phenomena

- We will now explore a set of relativistic phenomena which allow us to calculate some of these effects, in increasing order of difficulty ...

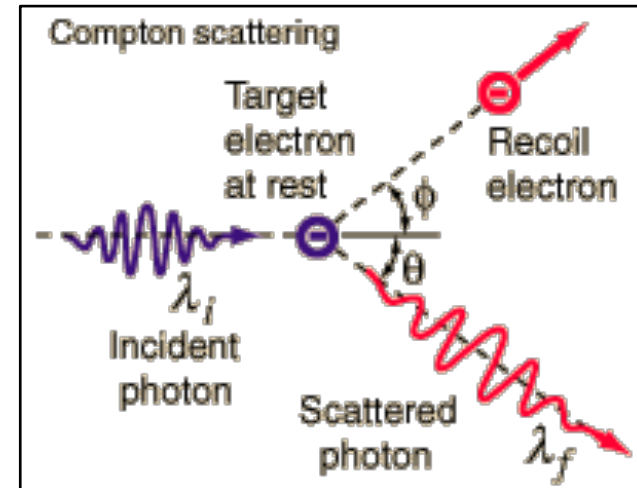
Nuclear energy



Pair production

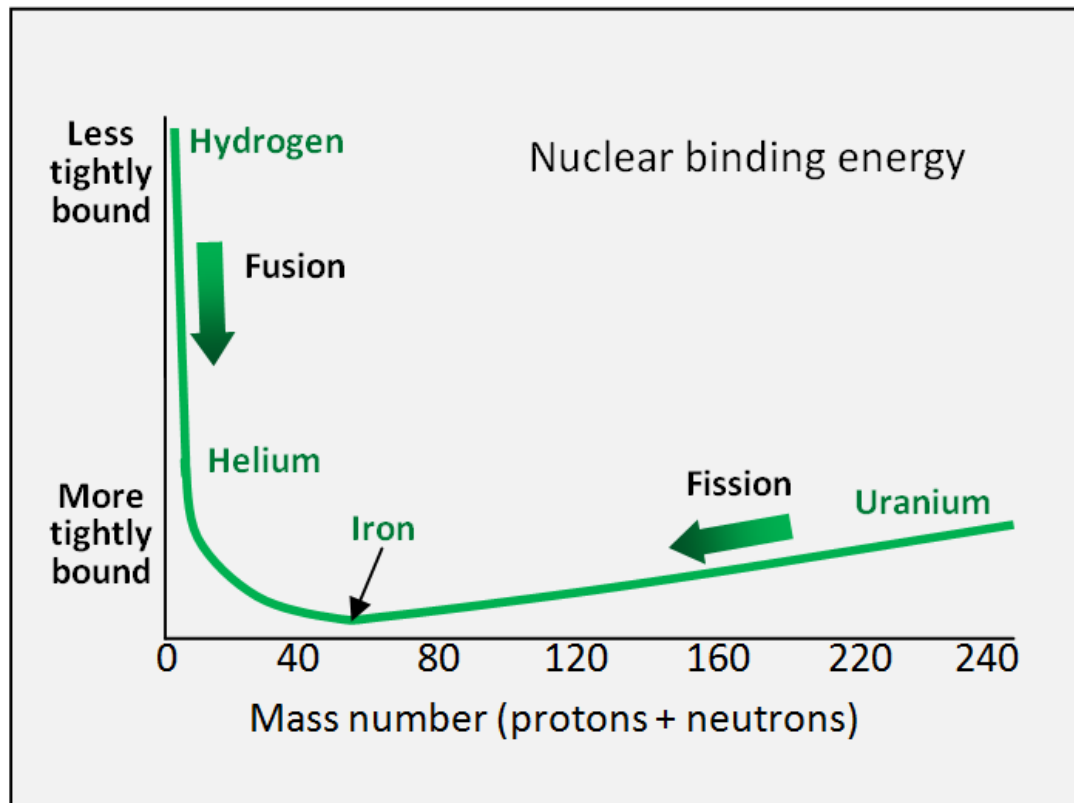


Compton effect



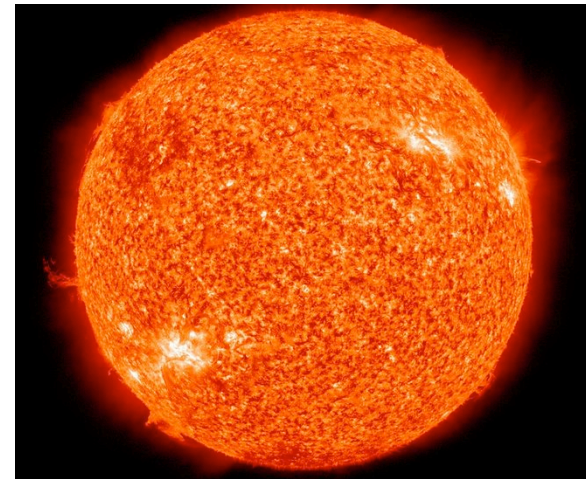
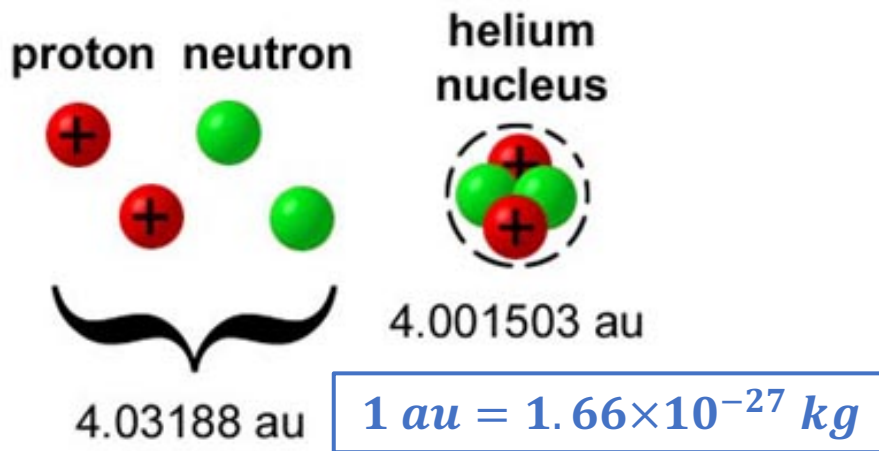
# Nuclear energy

- Energy may be released through the combination (**fusion**) or breaking-up (**fission**) of atomic nuclei



# Nuclear energy

- The amount of released energy (or, the binding energy) may be computed from **mass-energy equivalence**,  $E = m_0c^2$
- The Sun generates energy by **fusing hydrogen into helium**



[http://www.hk-phy.org/articles/fusion/fusion\\_e.html](http://www.hk-phy.org/articles/fusion/fusion_e.html)

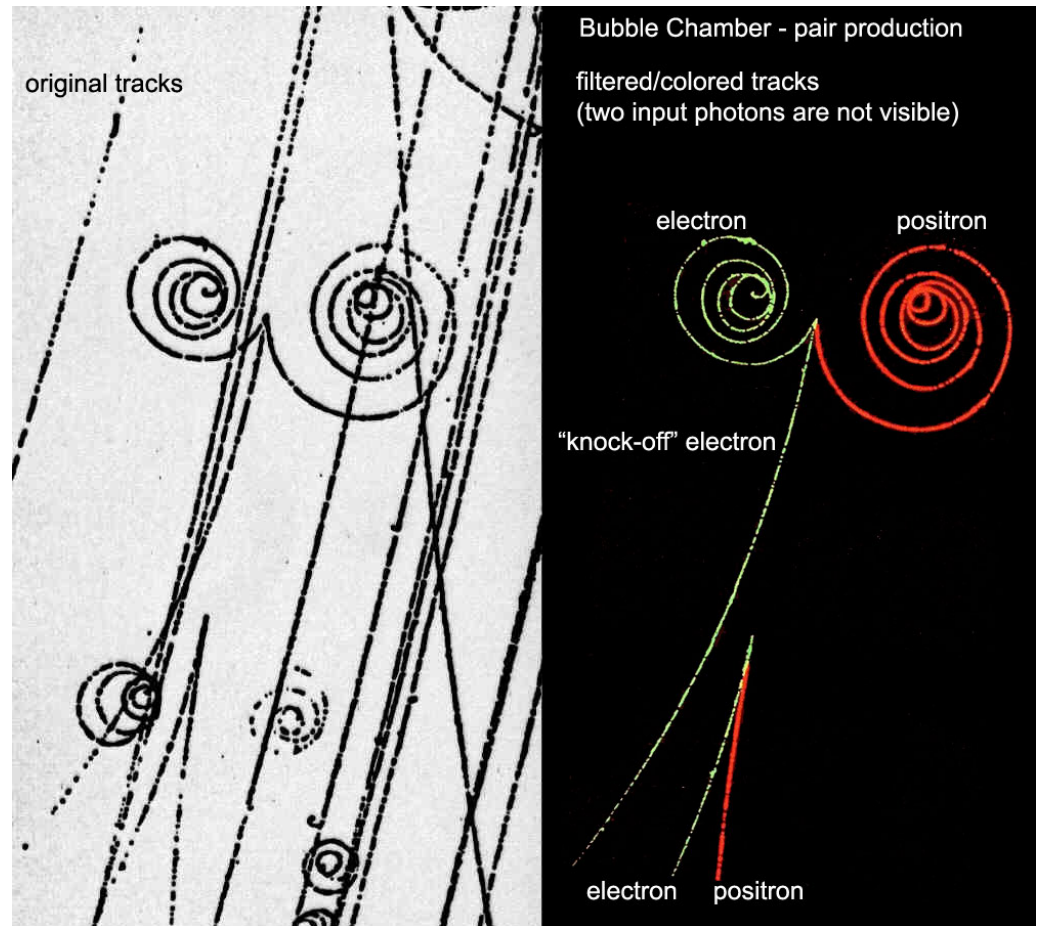
<https://www.space.com/19321-sun-formation.html>

- The luminosity of the Sun is  $3.8 \times 10^{26} \text{ W}$ . What mass of hydrogen is converted in the Sun every second?

# Pair production

- **Pair production** is the spontaneous creation of an electron and positron pair from electromagnetic radiation
- It's the main way that high-energy photons interact with matter
- It's a great example of the **interchangeability of matter and energy**

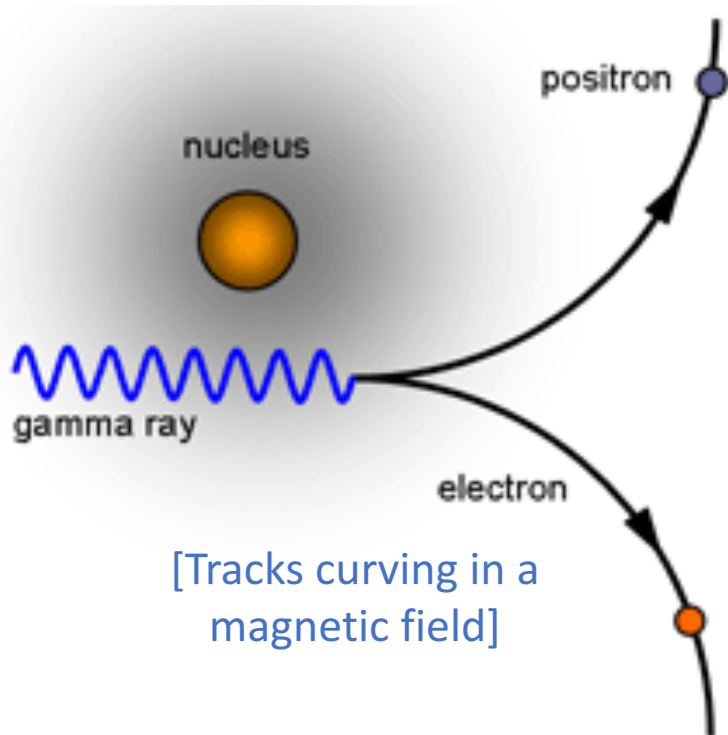
## Bubble chamber tracks



# Pair production

- Let's have a closer look at pair production [approximation!]

Mass of electron =  $9.1 \times 10^{-31} \text{ kg}$

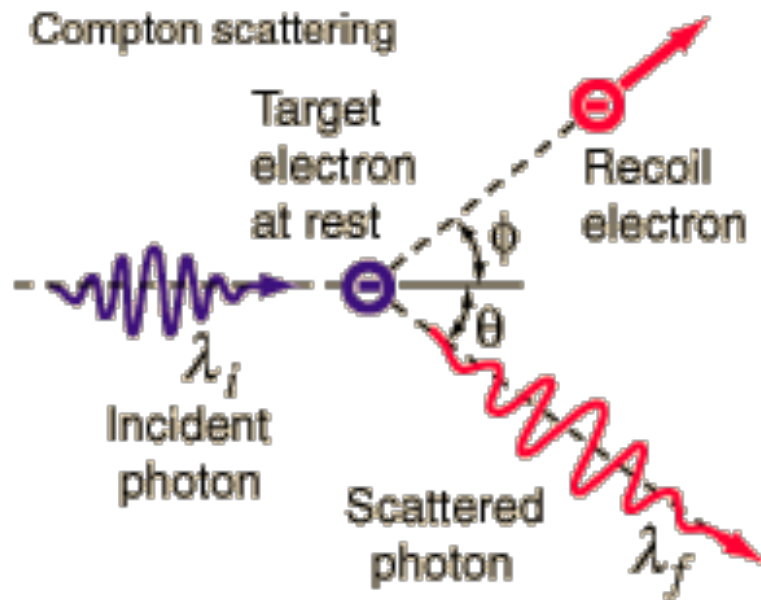


- Using mass-energy equivalence, what photon **threshold energy** is required to produce the electron/positron pair?
- What **wavelength** (and type) of radiation is required to achieve this?
- Why can this process only take place near an atomic nucleus? [Hint: consider conservation laws]

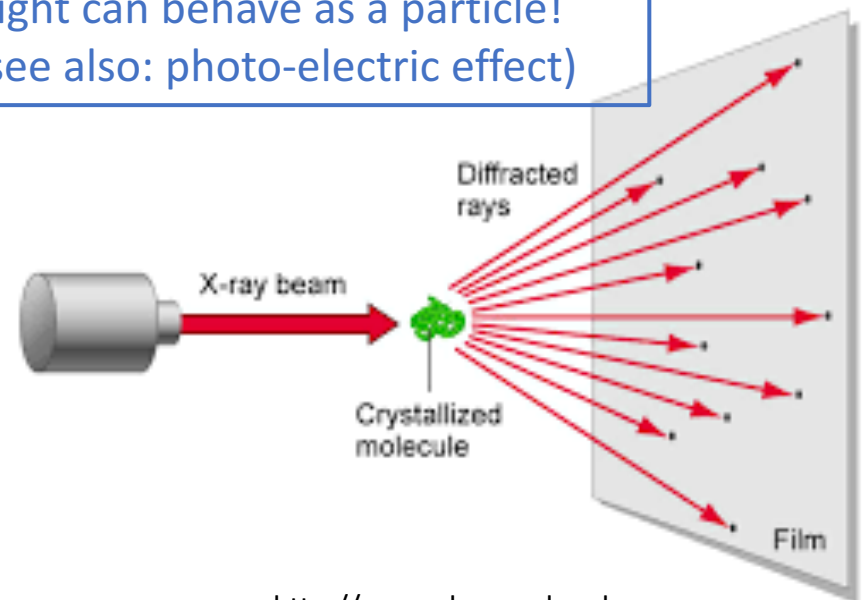


# Compton scattering

- The **Compton effect** is the scattering of a photon by a charged particle, resulting in a decrease in the energy (increase in the wavelength) of the radiation

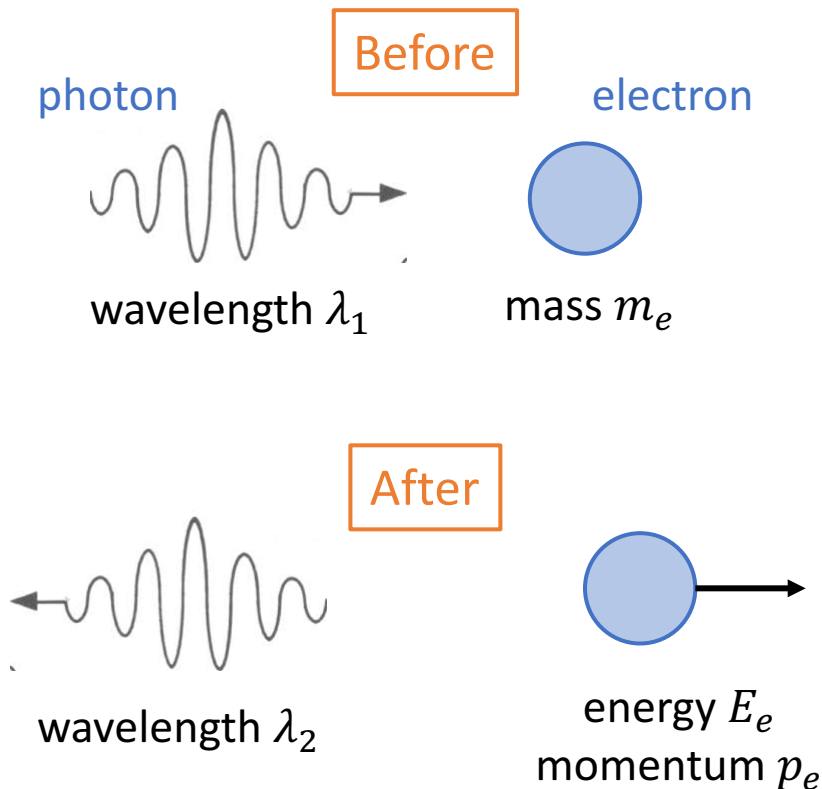


It's a beautiful demonstration that light can behave as a particle! (see also: photo-electric effect)



# Compton scattering

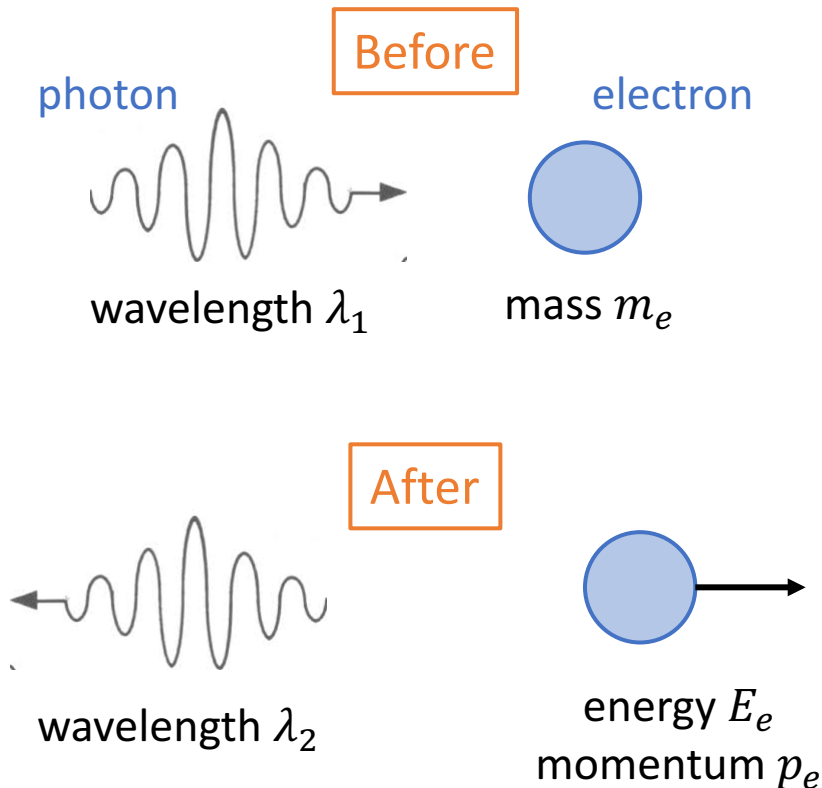
- Let's have a closer look at Compton scattering – how much does the wavelength change? [we'll consider a 1D case]



- Write an expression for the **momentum of the electron** after the collision,  $p_e$ , in terms of  $\lambda_1$  and  $\lambda_2$
- Write an expression for the **energy of the electron** after the collision,  $E_e$ , in terms of  $\lambda_1$ ,  $\lambda_2$  and  $m_e$
- Now use the fundamental relation  $E_e^2 - (p_e c)^2 = (m_e c^2)^2$ , to find an expression for  $\lambda_2 - \lambda_1$

# Compton scattering

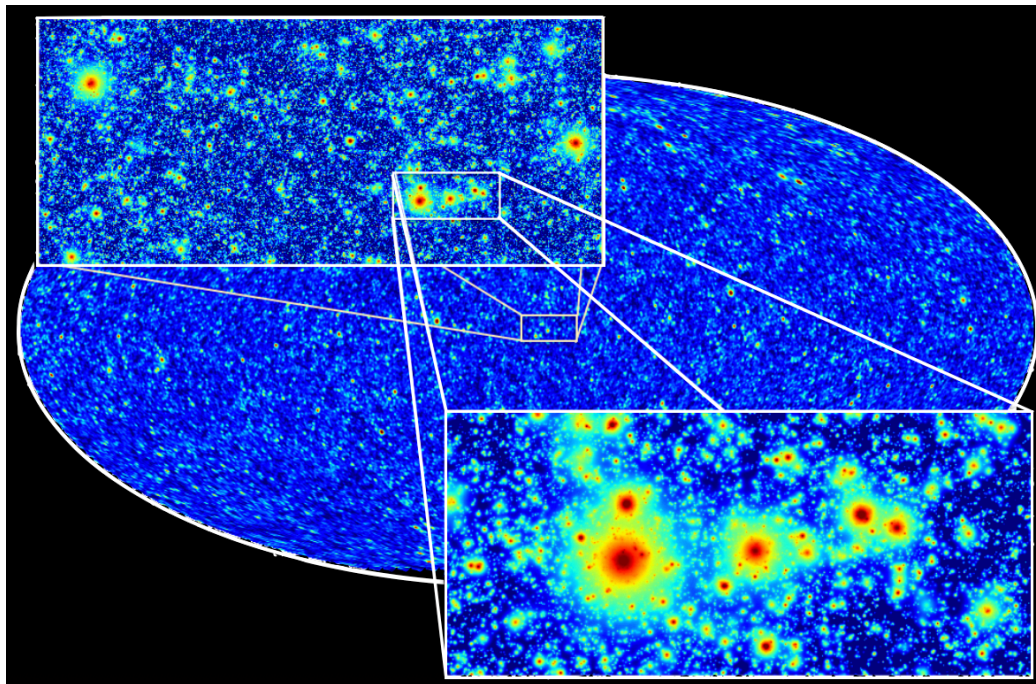
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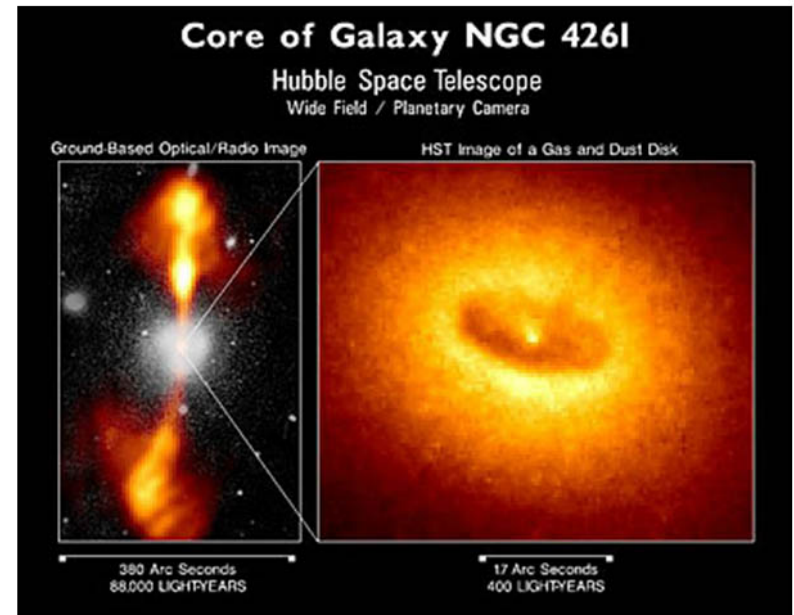
- Momentum of a photon is  $p = \frac{E}{c} = \frac{h}{\lambda'}$ , so  $\frac{h}{\lambda_1} = p_e - \frac{h}{\lambda_2}$ , or  $p_e = \frac{h}{\lambda_1} + \frac{h}{\lambda_2}$
- Energy before =  $\frac{hc}{\lambda_1} + m_e c^2$  and energy after =  $\frac{hc}{\lambda_2} + E_e$ , so we find  $E_e = \left(\frac{hc}{\lambda_1} + m_e c^2\right) - \frac{hc}{\lambda_2}$
- Substitute these two expressions in  $E_e^2 - (p_e c)^2 = (m_e c^2)^2$  and cancel some terms, we find:  $\lambda_2 - \lambda_1 = \frac{2h}{m_e c}$

# Compton scattering

- **Inverse Compton scattering** – where high-energy electrons impart energy to photons – also occurs in astrophysics, in *clusters of galaxies* and *Active Galactic Nuclei*



<http://cita.utoronto.ca/~malvarez/research/>

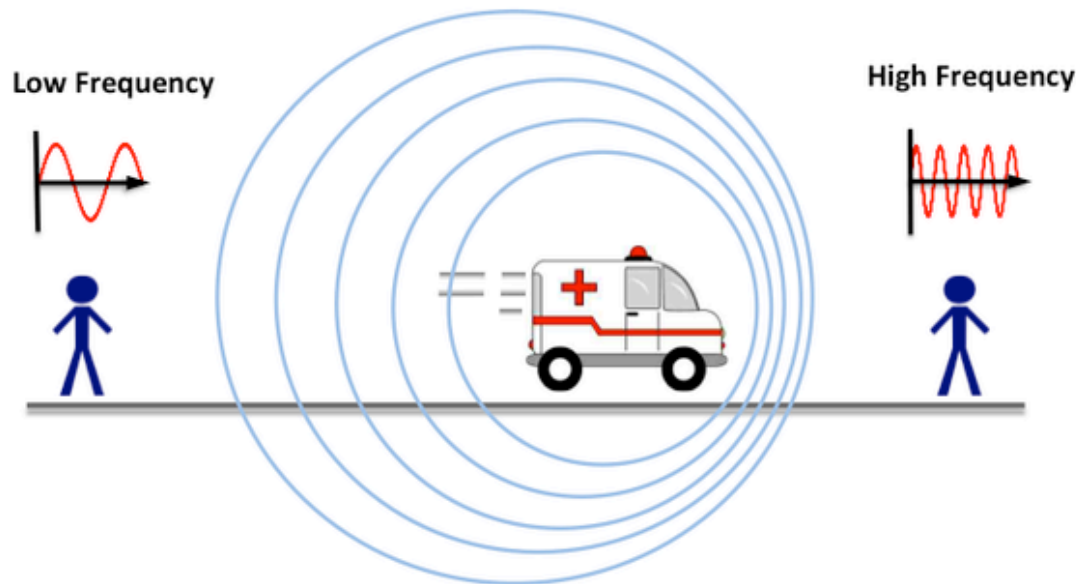


**Figure 5:** Active galaxy NGC 4261 at radio and optical wavelengths.

<http://www.jeffstanger.net/Astronomy/emissionprocesses.html>

# The Doppler effect

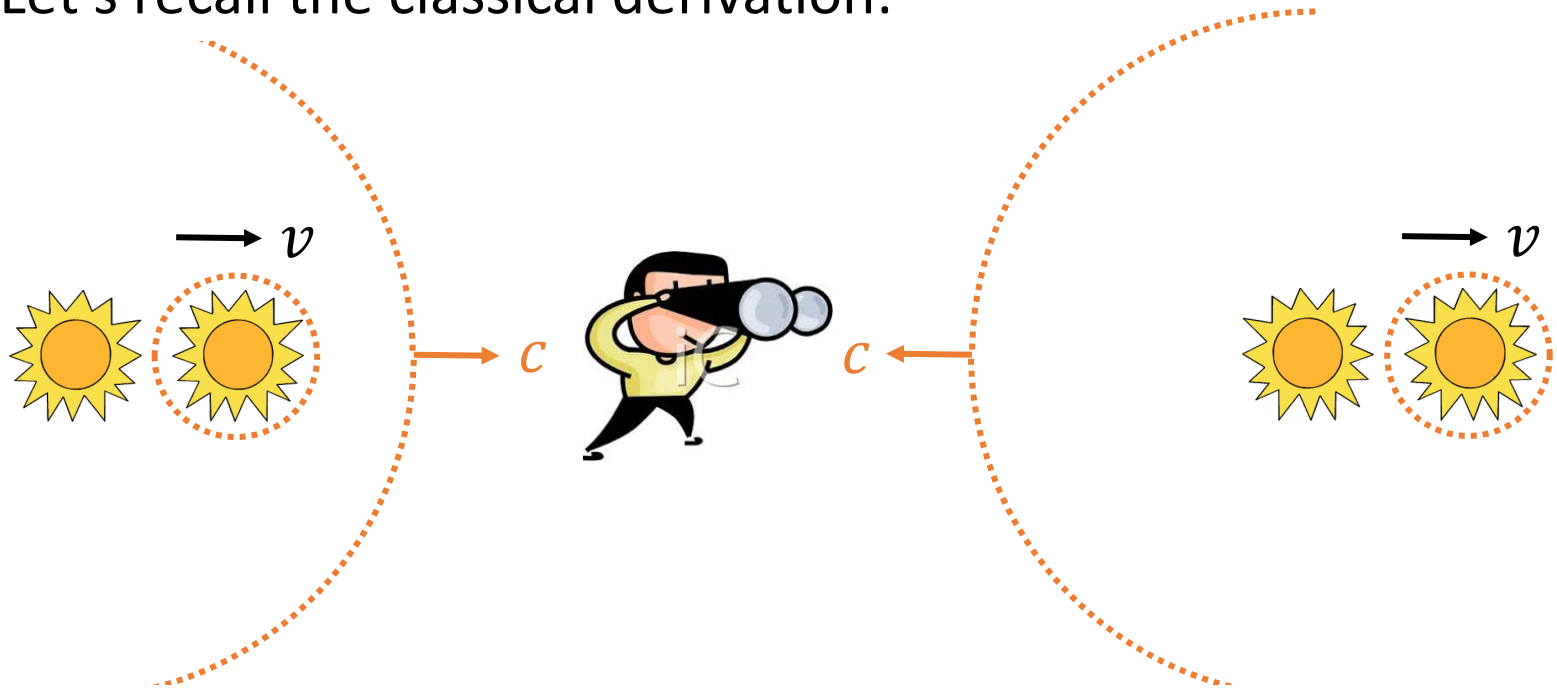
- The **Doppler effect** describes the change in frequency of a wave due to the relative motion of source and observer



- Since times transform differently in relativity, **the Doppler effect is modified at high speeds**

# The Doppler effect

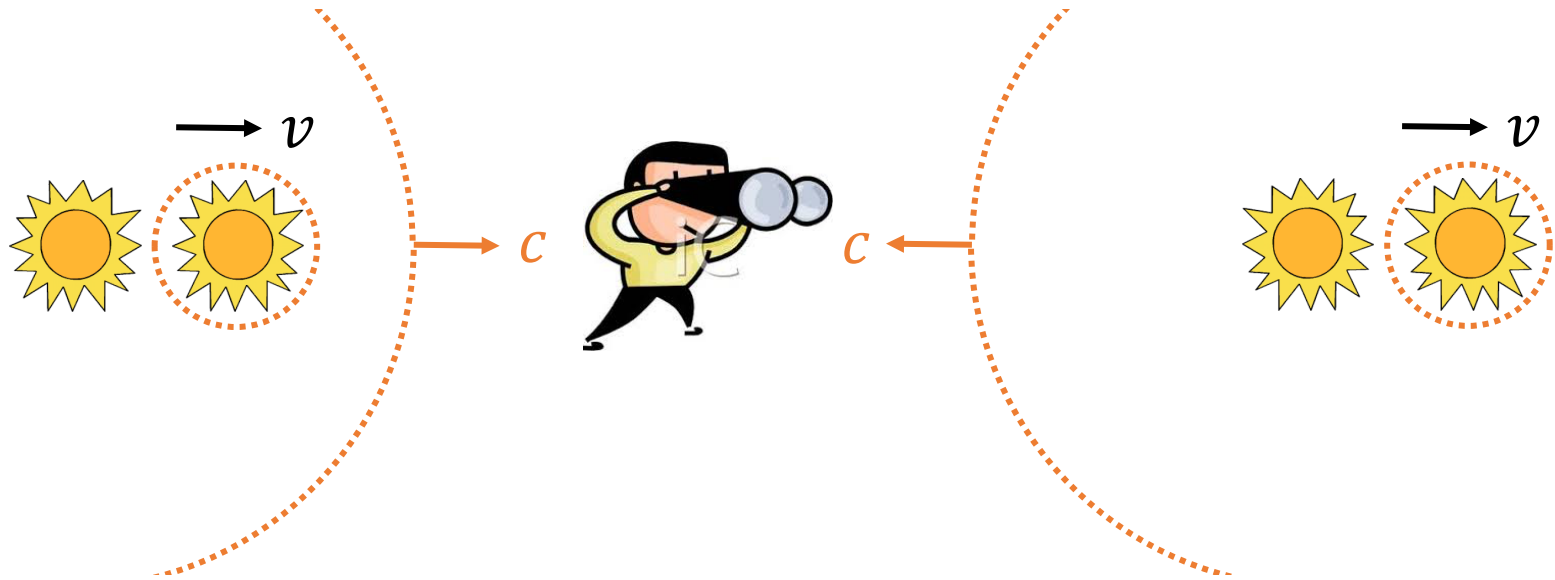
- Let's recall the classical derivation:



- Source pulses every time  $\Delta t$  (frequency  $f_e = 1/\Delta t$ , wavelength  $\lambda_e = c/f_e$ )
- In this time, the wave travels distance  $c \Delta t$  and the source travels  $v \Delta t$
- Wavelength measured by observer  $\lambda_o = c \Delta t \pm v \Delta t = \lambda_e (1 \pm v/c)$

# The Doppler effect

- In relativity, there is additional **time dilation** such that the time interval between pulses in the observer frame is  $\gamma \Delta t$



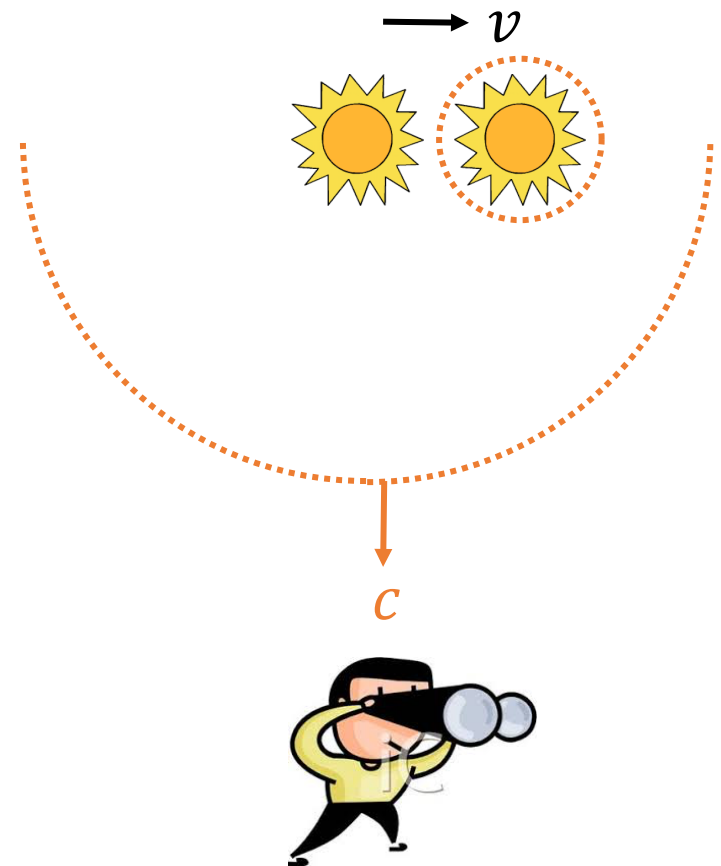
- Wavelength measured by observer  $\lambda_o = c \gamma \Delta t \pm v \gamma \Delta t = \lambda_e \frac{1 \pm v/c}{\sqrt{1 - v^2/c^2}}$

$$\lambda_e = \lambda_o \sqrt{\frac{1 + v/c}{1 - v/c}} \quad (\text{receding}) \quad \lambda_e = \lambda_o \sqrt{\frac{1 - v/c}{1 + v/c}} \quad (\text{approaching})$$

# The Doppler effect

- In relativity, time dilation causes a **Doppler effect for transverse motion** which doesn't exist in classical physics!

- $\lambda_e = \lambda_o \frac{1}{\sqrt{1-v^2/c^2}}$  (transverse)





# The Doppler effect

- A relativity student is caught running a red light on the rocket freeway. In court they plead they were driving so fast that the red light ( $\lambda = 650 \text{ nm}$ ) looked green ( $\lambda = 530 \text{ nm}$ ). How fast would their rocket car have been travelling?



[Problem 8-20 from  
Taylor & Wheeler,  
*Spacetime Physics*]

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