#### **Class 10: Einstein Equation**

In this class we will introduce Einstein's equations of General Relativity, which provide the all-important link between space-time curvature and matter-energy

# **Class 10: Einstein Equation**

At the end of this session you should be able to ...

- ... recognize **Einstein's equations of General Relativity**, which link space-time curvature to matter-energy
- ... be familiar with the components of the Einstein equations the **Ricci tensor**  $R_{\mu\nu}$  and the **Ricci scalar** R
- ... show that Einstein's equations recover Newton's laws in the appropriate weak-field limit
- ... understand how, mathematically, we can test whether or not we are in empty space, free of matter

# Linking curvature and matter

#### John Wheeler has a famous summary of General Relativity:



Spacetime tells matter how to move; matter tells spacetime how to curve.

- John Archibald Wheeler -

#### AZQUOTES

# Linking curvature and matter

Space-time tells matter how to move (geodesic equation) Matter tells space-time how to curve (Einstein's equation)



https://www.quora.com/What-is-the-best-way-to-imagine-space-time

### Ricci tensor

- We have seen that general space-time curvature is described by the **Riemann tensor**  $R_{\mu\lambda\nu}^{\kappa}$  (4 indices!)
- However, the metric of space-time, and its energymomentum content, are described by the simpler objects  $g_{\mu\nu}$  and  $T_{\mu\nu}$  (2 indices!)
- We can produce an object of corresponding complexity from the Riemann tensor by contracting it as follows:  $R_{\mu\nu} = R_{\mu\lambda\nu}^{\lambda}$
- This is known as the **Ricci tensor**,  $R_{\mu\nu}$  it is an "average of the Riemann tensor"
- Einstein proposed his famous equation linking  $R_{\mu\nu}$  and  $T_{\mu\nu}$

• Einstein proposed that:

 $\frac{8\pi G}{{}_{\mathcal{A}}}T_{\mu\nu}$  $G_{\mu\nu}$ 



https://www.nobelprize.org/nobel\_prizes/ physics/laureates/1921/einstein-bio.html

- $G_{\mu\nu}$  is called the **Einstein tensor**
- It is related to the Ricci tensor  $R_{\mu\nu}$  by  $G_{\mu\nu} = R_{\mu\nu} \frac{1}{2}Rg_{\mu\nu}$
- *R* is the **Ricci scalar**,  $R = R^{\mu}_{\mu} = g^{\mu\nu}R_{\mu\nu}$

Matter tells spacetime how to curve (Einstein's equation)

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

- This is "10 equations in one"! ( $\mu$  and  $\nu$  can each take on 4 values, but the equation is symmetric,  $T_{\mu\nu} = T_{\nu\mu}$ )
- Each equation is a non-linear differential equation for  $g_{\mu\nu}$



- Why Einstein's equation?? First, it is mathematically consistent
- Second, it reproduces Newtonian gravity in the weak-field limit
- Third, it makes predictions which have been experimentally verified – e.g. the perihelion of Mercury, gravitational deflection



http://www.eniscuola.net/en/mediateca/precession-of-mercury/



http://andybohn.com/research/lensing.html

- The Einstein equation is  $R_{\mu\nu} \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$
- If we multiply both sides by  $g^{\lambda\mu}$  and perform a little algebra, we find a useful re-write:  $R_{\mu\nu} = \frac{8\pi G}{c^4} \left( T_{\mu\nu} \frac{1}{2} g_{\mu\nu} T_{\lambda}^{\lambda} \right)$
- We can also show after some algebra that ...

$$R_{\mu\nu} = \partial_{\lambda}\Gamma^{\lambda}_{\mu\nu} - \partial_{\nu}\Gamma^{\lambda}_{\mu\lambda} + \Gamma^{\kappa}_{\kappa\lambda}\Gamma^{\lambda}_{\mu\nu} - \Gamma^{\lambda}_{\nu\kappa}\Gamma^{\kappa}_{\mu\lambda}$$

 There's no need to memorize this! But the point is that the Christoffel symbols (which determine the geodesics), and the Ricci tensor (which links to the matter-energy), are all completely determined by the metric (after heavy calculating!)

# What is "empty space"?

• How can we tell that we are in empty space (free of matter)?



- It would wrong to say that  $g_{\mu\nu} = 0$  in empty space, of course
- $\Gamma_{\mu\nu}^{\lambda} \neq 0$  empty space can be curved by the matter outside it!
- Einstein's equations show us that  $R_{\mu\nu} = 0$  in empty space
- For example, to verify the Schwarzschild metric we used in the previous class, we would need to show that  $R_{\mu\nu} = 0$