

Class 6: Curved Space and Metrics

In this class we will discuss the meaning of spatial curvature, how distances in a curved space can be measured using a metric, and how this is connected to gravity

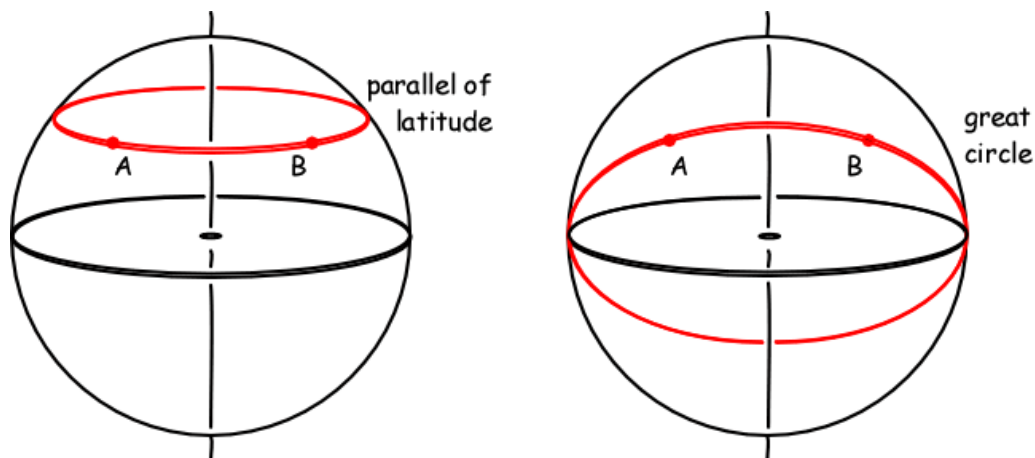
Class 6: Curved Space and Metrics

At the end of this session you should be able to ...

- ... describe the **geometrical properties of curved spaces** compared to flat spaces, and how observers can determine whether or not their space is curved
- ... know how the **metric of a space** is defined, and how the metric can be used to compute distances and areas
- ... make the connection between **space-time curvature and gravity**
- ... apply the **space-time metric** for Special Relativity and General Relativity to determine proper times and distances

Properties of curved spaces

- In the last Class we discussed that, according to the Equivalence Principle, objects “*move in straight lines in a curved space-time*”, in the presence of a gravitational field
- So, what is a straight line on a curved surface? We can define it as *the shortest distance between 2 points*, which mathematicians call a **geodesic**

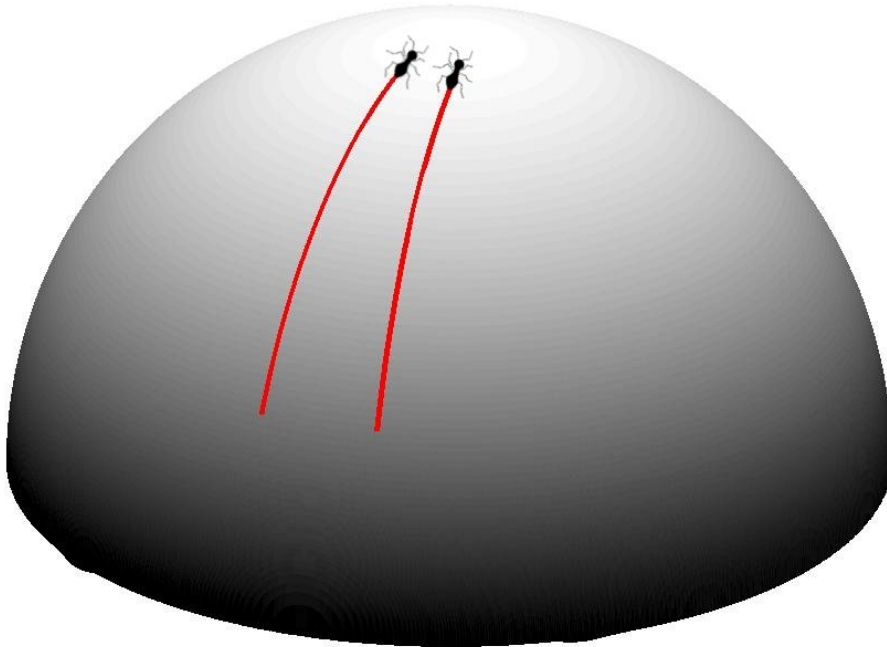


https://www.pitt.edu/~jdnorton/teaching/HPS_0410/chapters/non_Euclid_curved/index.html

<https://www.quora.com/What-are-the-reasons-that-flight-paths-especially-for-long-haul-flights-are-seen-as-curves-rather-than-straight-lines-on-a-screen-ls-map-distortion-the-only-reason-Or-do-flight-paths-consider-the-rotation-of-the-Earth>

Properties of curved spaces

Equivalently, a geodesic is a path that would be travelled by an ant walking straight ahead on the surface!



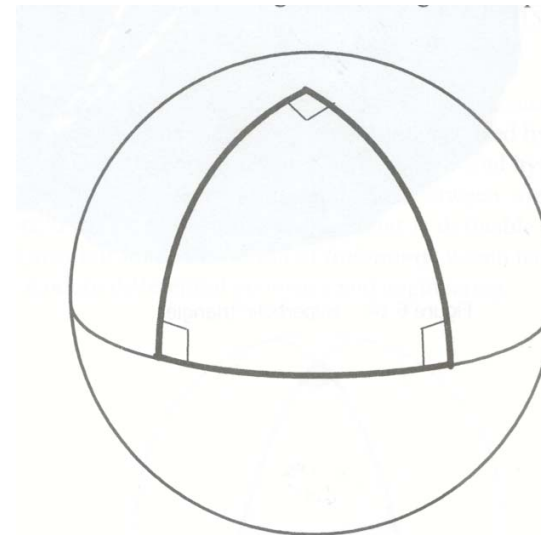
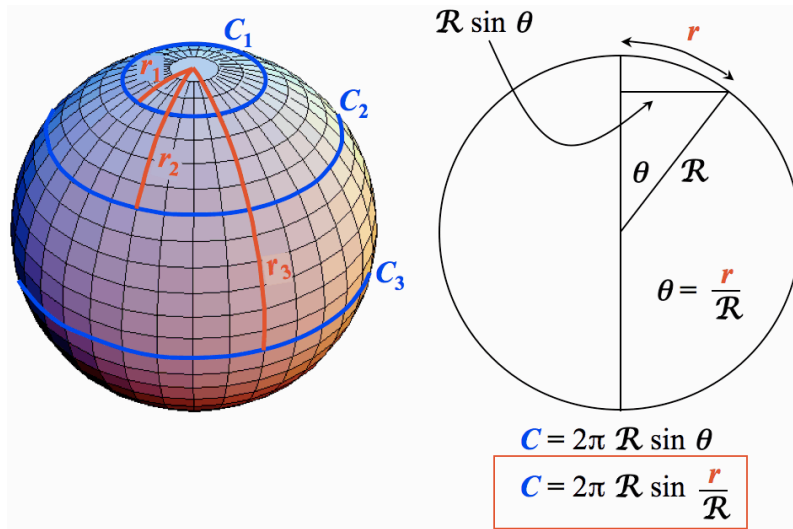
<http://astronomy.nmsu.edu/geas/lectures/lecture28/slide03.html>

- Consider two ants starting from different points on the Equator, both walking North
- These geodesics are both “straight lines”, but they are converging
- **Parallel lines converge or diverge on a curved surface**

Properties of curved spaces

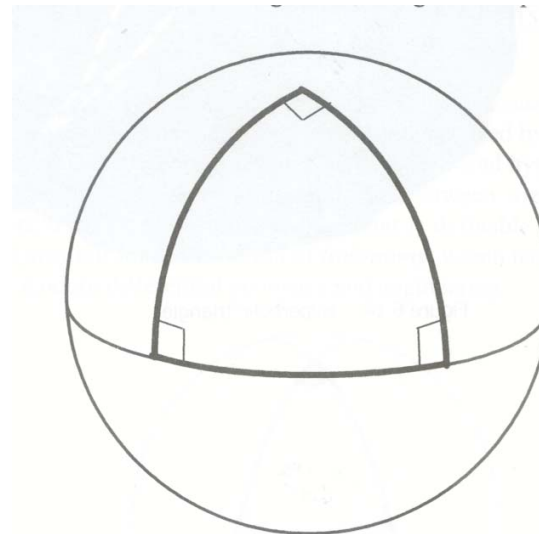
Some other counter-intuitive properties of curved surfaces:

- The circumference of a circle of radius r is not $2\pi r$
- The area of a circle of radius r is not πr^2
- The angles of a triangle do not add up to 180°



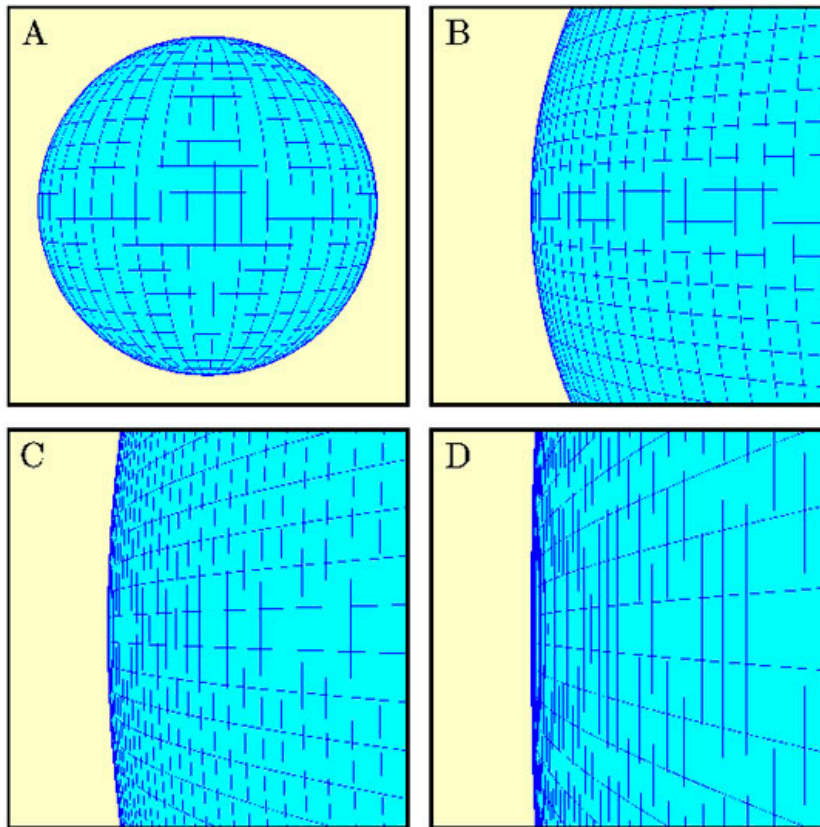
Properties of curved spaces

- It's easy to visualize curvature by thinking of a *2D curved surface embedded in a 3D Euclidean space*
- However, **curvature is intrinsic to a surface** and can be determined without external reference – Earth dwellers can know the Earth is curved without seeing it from space!



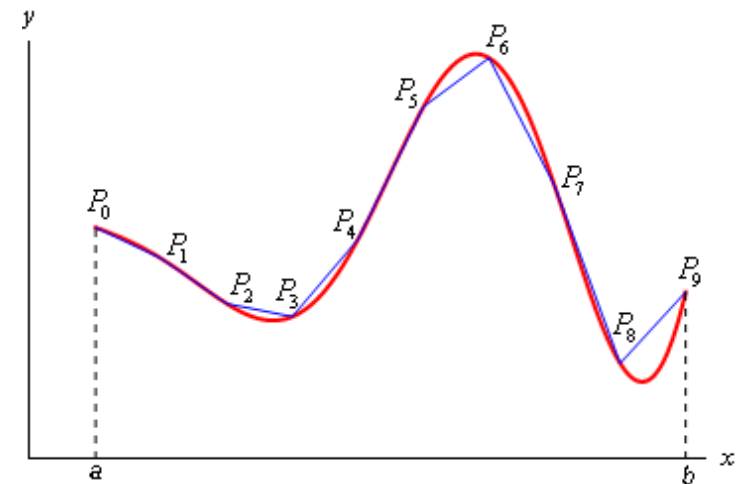
Properties of curved spaces

- Zooming into a small region, a **curved surface is locally flat** (just as a page of an atlas represents a piece of the globe)



<https://ned.ipac.caltech.edu/level5/March05/Guth/Guth1.html>

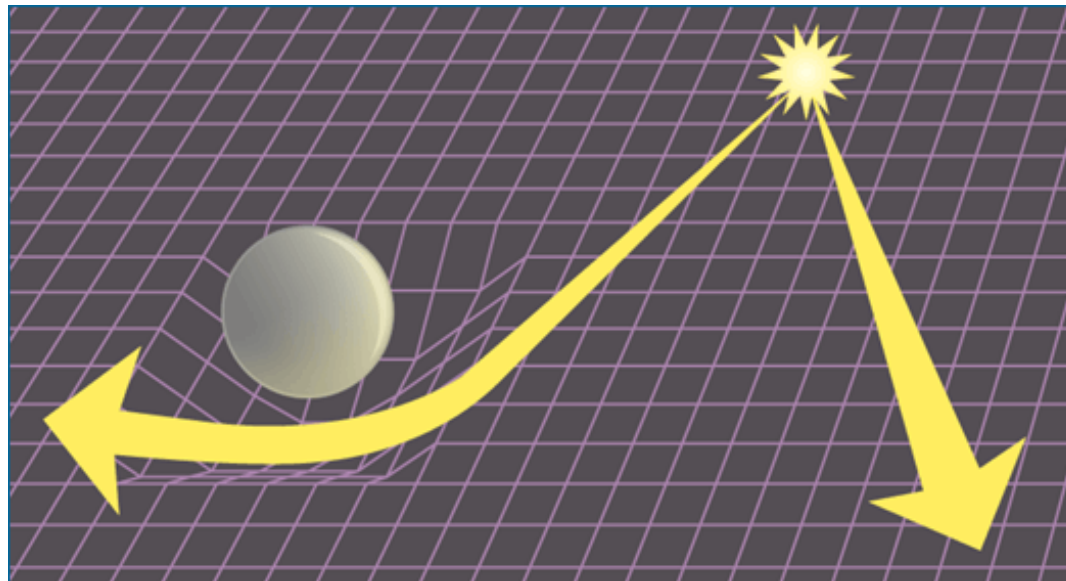
It's analogous to calculus, where we build curves out of straight lines ...



<http://tutorial.math.lamar.edu/Classes/CalcII/ArcLength.aspx>

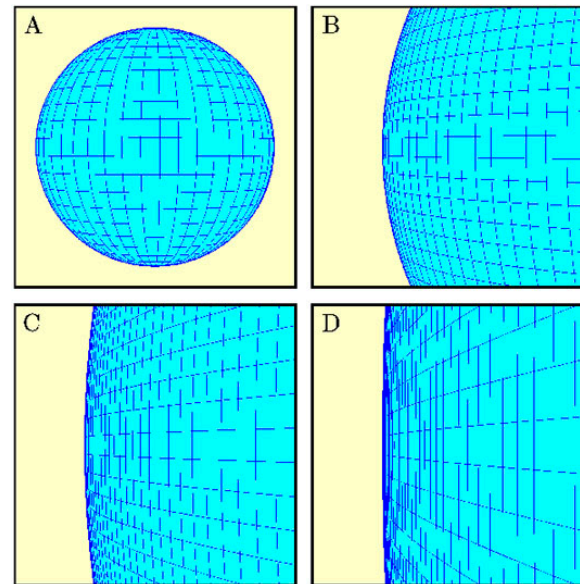
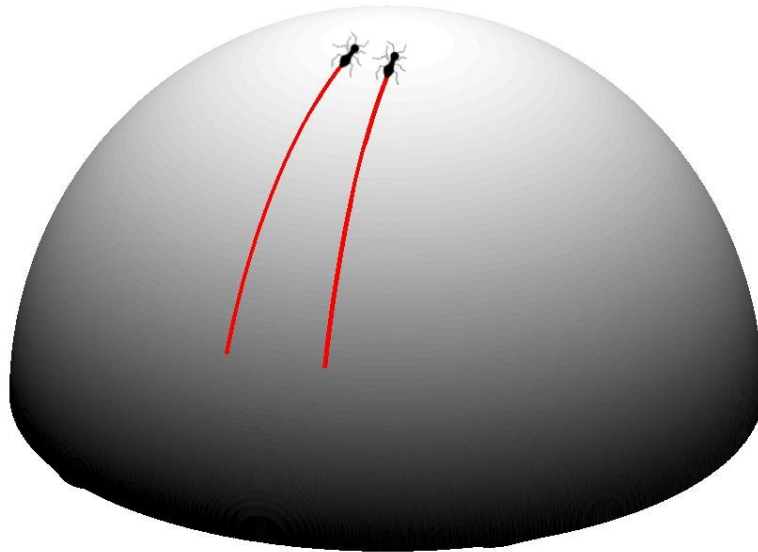
Curvature and gravity

- *What has this got to do with gravity?*
- Gravity can be represented as the **curvature of space-time**
- **Objects travel on a geodesic** in the curved space-time, that *extremizes the space-time interval* between the two points



Curvature and gravity

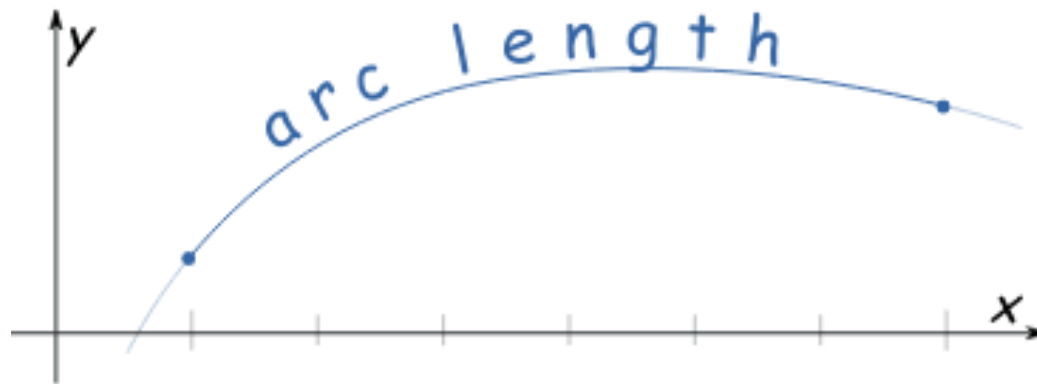
- Other forces **cause particles to deviate from geodesics** – e.g., *if 2 ants on a curved surface are connected by a solid bar, the force would push them off their geodesics*



- Just as a curved surface is locally flat, a **curved space-time can locally be described by an inertial frame** (of a freely-falling observer), but there is no extended inertial frame

The metric of a space

- How do we measure lengths and angles in a Cartesian space with co-ordinates (x, y) ?



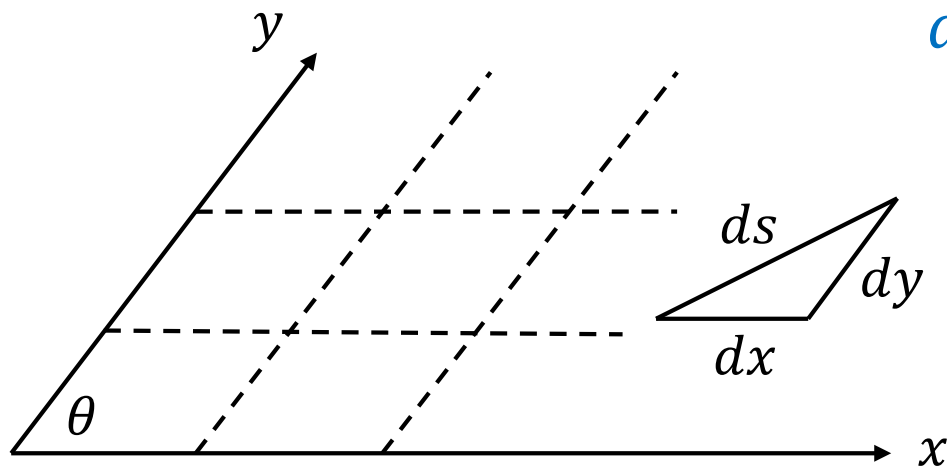
- Breaking the arc into small pieces: $ds^2 = dx^2 + dy^2$

$$s = \int_A^B \sqrt{\left(\frac{dx}{d\lambda}\right)^2 + \left(\frac{dy}{d\lambda}\right)^2} d\lambda$$

λ parameterizes the curve [i.e., $x(\lambda), y(\lambda)$]

The metric of a space

- What happens in a *tilted co-ordinate system*?



$$ds^2 = dx^2 + dy^2 - 2 dx dy \cos \theta$$

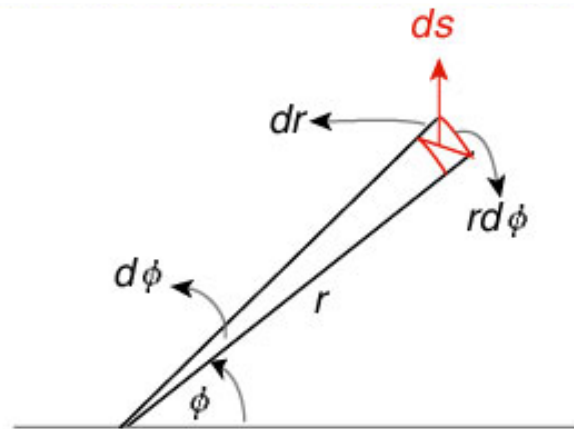
$$= \sum_i \sum_j g_{ij} dx^i dx^j$$

where $i = \{1,2\}$, $x^{(1)} = x$, $x^{(2)} = y$

- Using the summation convention, this can be written as $ds^2 = g_{ij} dx^i dx^j$ where g_{ij} is the metric of the space
- In this case, $g_{11} = 1$, $g_{22} = 1$, $g_{12} = g_{21} = -\cos \theta$

The metric of a space

- This example shows that *the metric determines the geometry, but the geometry does not determine the metric*
- For a given geometry, we can generate many possible metrics through co-ordinate transformations



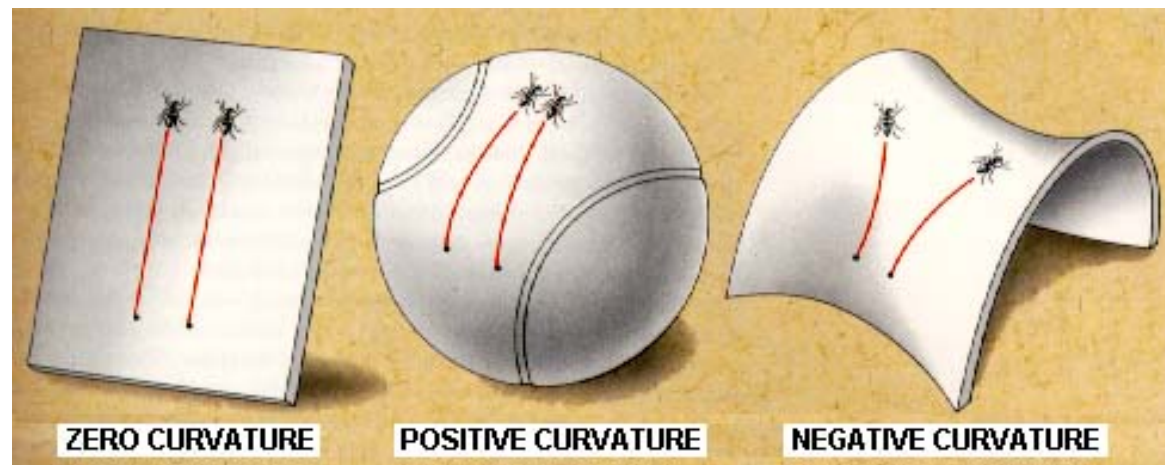
For example – Cartesian and polar co-ordinates

$$(ds)^2 = (dx)^2 + (dy)^2 = (dr)^2 + r^2(d\phi)^2$$

- **The metric encodes information about both the geometry and the co-ordinate system**

How do we define curvature?

- How can observers on a surface **quantify the amount of curvature at a point?**



<https://starchild.gsfc.nasa.gov/docs/StarChild/questions/question35.html>

- Move a geodesic distance ε in all directions to form a “circle” in the space and measure its area A
- The curvature at a point is then $\lim_{\varepsilon \rightarrow 0} \left[\frac{12}{\varepsilon^2} \left(1 - \frac{A}{\pi \varepsilon^2} \right) \right]$

The space-time metric

- The metric of space-time tell us **how to measure the space-time interval between events** (i.e., proper times/distances)

- *Special Relativity*: $ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$ where $\eta_{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$

- Note that *this geometry is not Euclidean* because $\eta_{00} = -1$ (the locus of events separated by constant ds is a hyperbola, not a circle) – it is known as a “**Minkowski geometry**”
- In *General Relativity*, in a frame containing a gravitational field with curved space-time, $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$ where $g_{\mu\nu}$ is the **space-time metric** – a more complicated function of x^μ

The space-time metric

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

- What is the structure of this function $g_{\mu\nu}$? It's a matrix, where each element is a function of the co-ordinates x^μ
- μ and ν run over 4 indices but, from the above equation, the metric must be symmetric ($g_{\mu\nu} = g_{\nu\mu}$), so there are **10 functions in general** – *this is why GR is complicated!!*

The space-time metric

- Consider a clock at rest in the Earth's frame, which ticks every dt seconds. Is this a proper time interval $d\tau$?
- **No**, because the clock is not in an inertial frame (it is not freely falling)
- The space-time interval between the ticks is $ds^2 = g_{00}(x^i) (c dt)^2$, since $dx^i = 0$
- Since $ds^2 = -c^2 d\tau^2$, the proper time interval is $d\tau = \sqrt{-g_{00}(x^i)} dt$
- **Time runs differently at each point of a gravitational field**

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$



The space-time metric

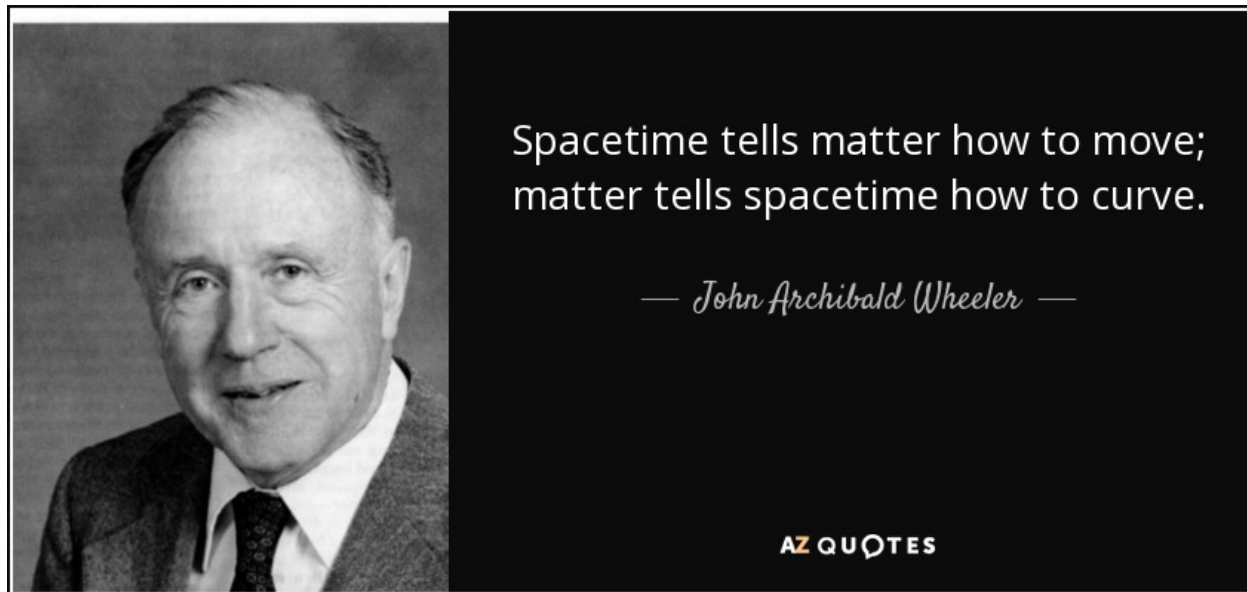
- A concrete example is a weak gravitational field with potential $\phi(\vec{x})$
- We will show later in the course that:
$$g_{00}(\vec{x}) = -1 - 2\phi/c^2$$
- At height h in a simple vertical gravitational field, $\phi = gh$
- So the ticking period of the clock varies as $dt = d\tau/\sqrt{1 + 2gh/c^2}$

$$d\tau = \sqrt{-g_{00}(\vec{x})} dt$$



What determines the metric?

- The question “*what determines the space-time metric*” is the same question as – “*what generates gravity*”. The answer is **the distribution of matter and energy**



- Later in the course, we will study the equation which links the space-time curvature to the distribution of matter