



Multi-Object Adaptive Optics
(MOAO)

Demonstration Instrument
Subaru Telescope

Richard West



University
of Victoria



TOHOKU
UNIVERSITY



Adaptive Optics
Lab

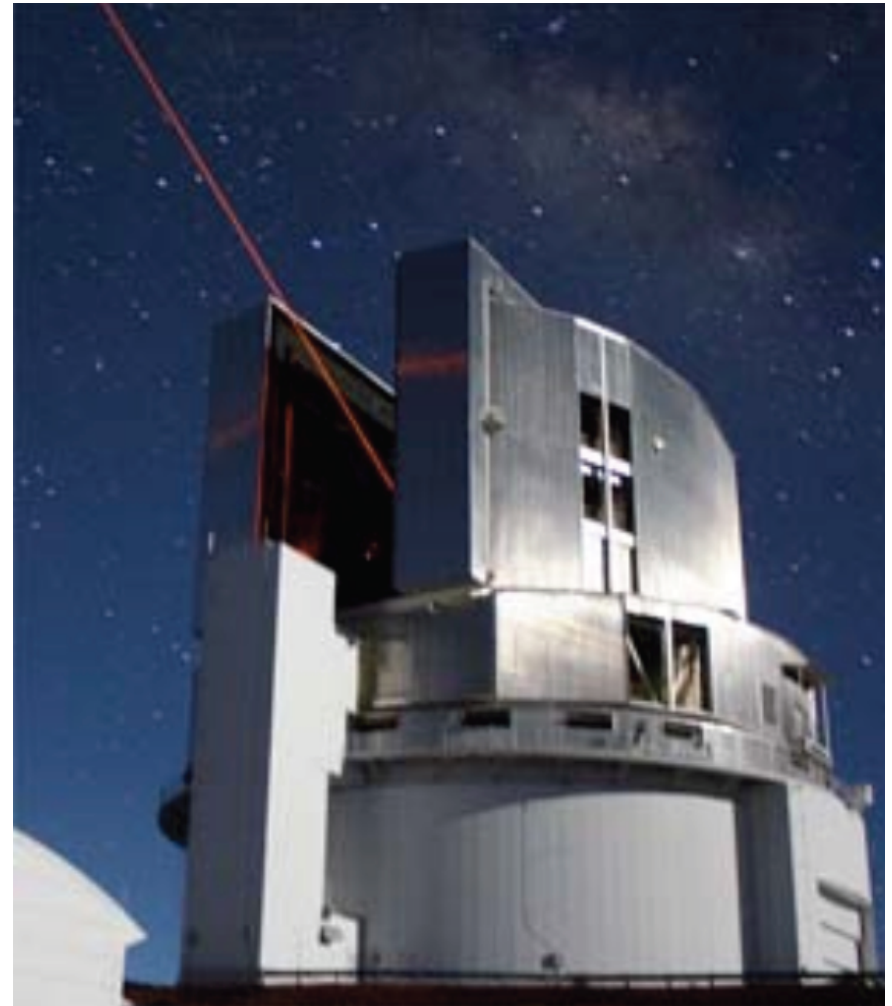
Herzberg
Institute

Subaru
Observatory

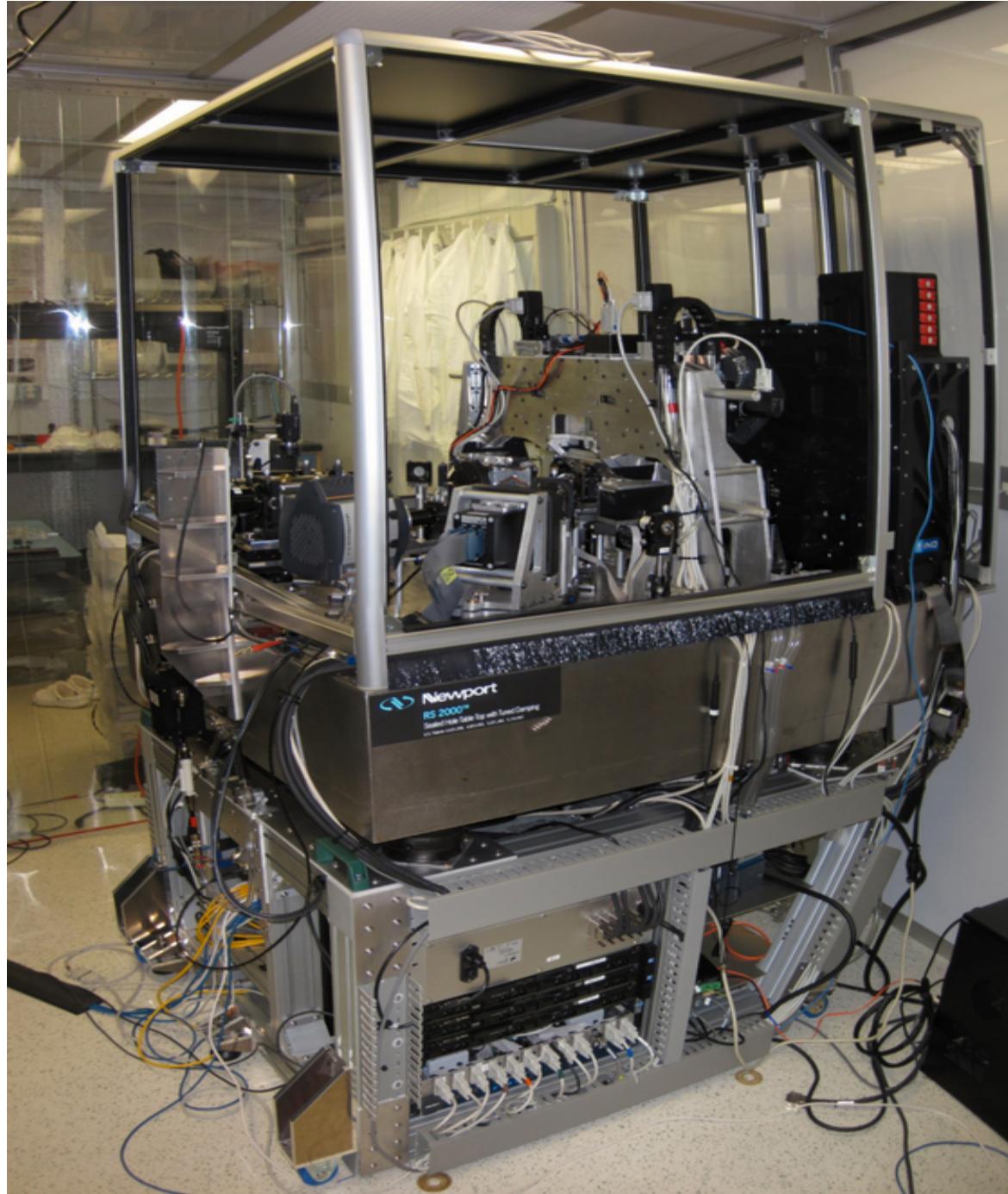
Astronomical
Institute

Institut National
d'Optique

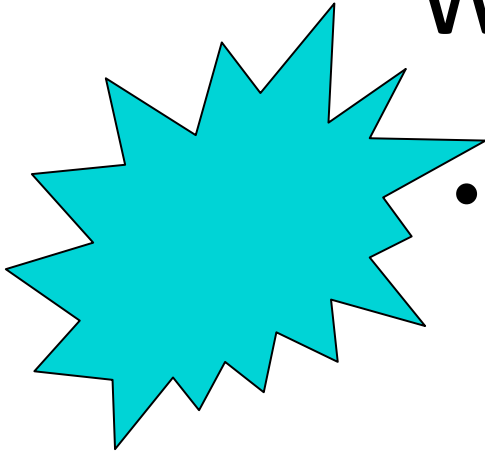
- Introduction
- MOAO demonstration on Subaru 2014/2015.
- AO review.
- Why build this demonstrator?
- What's next?



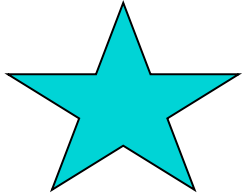
RAVEN



What's in an image



- Image Blurred by atmospheric effects.



- Image corrected by a TT mirror

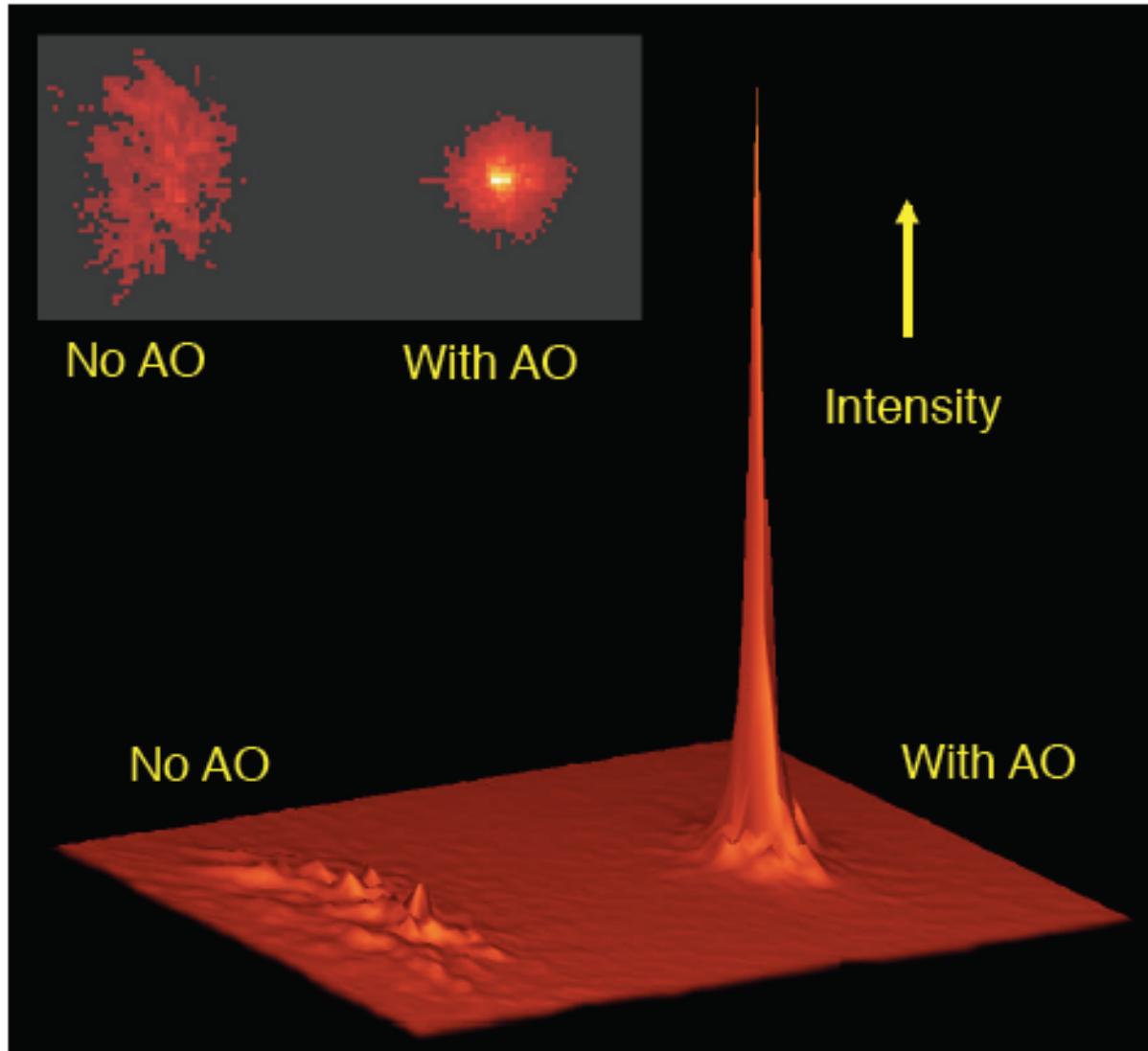


- Image further corrected by a DM



- Equivalent HST image

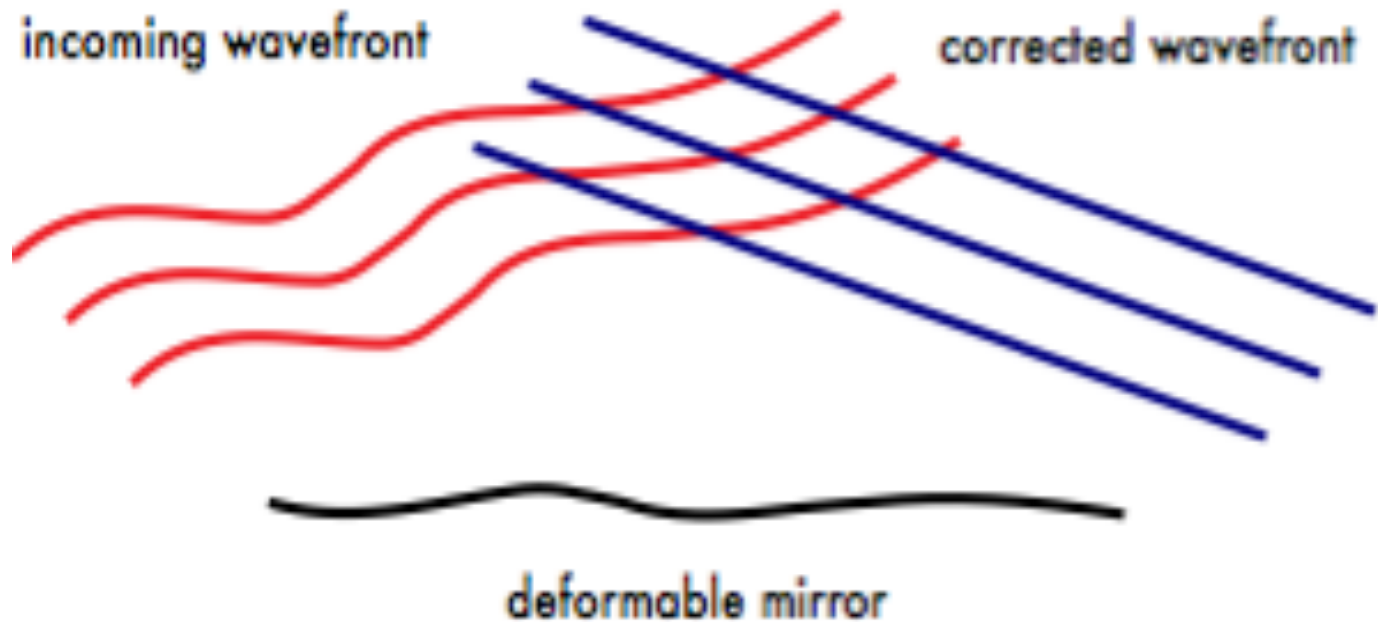
Adaptive optics increases peak intensity of a point source



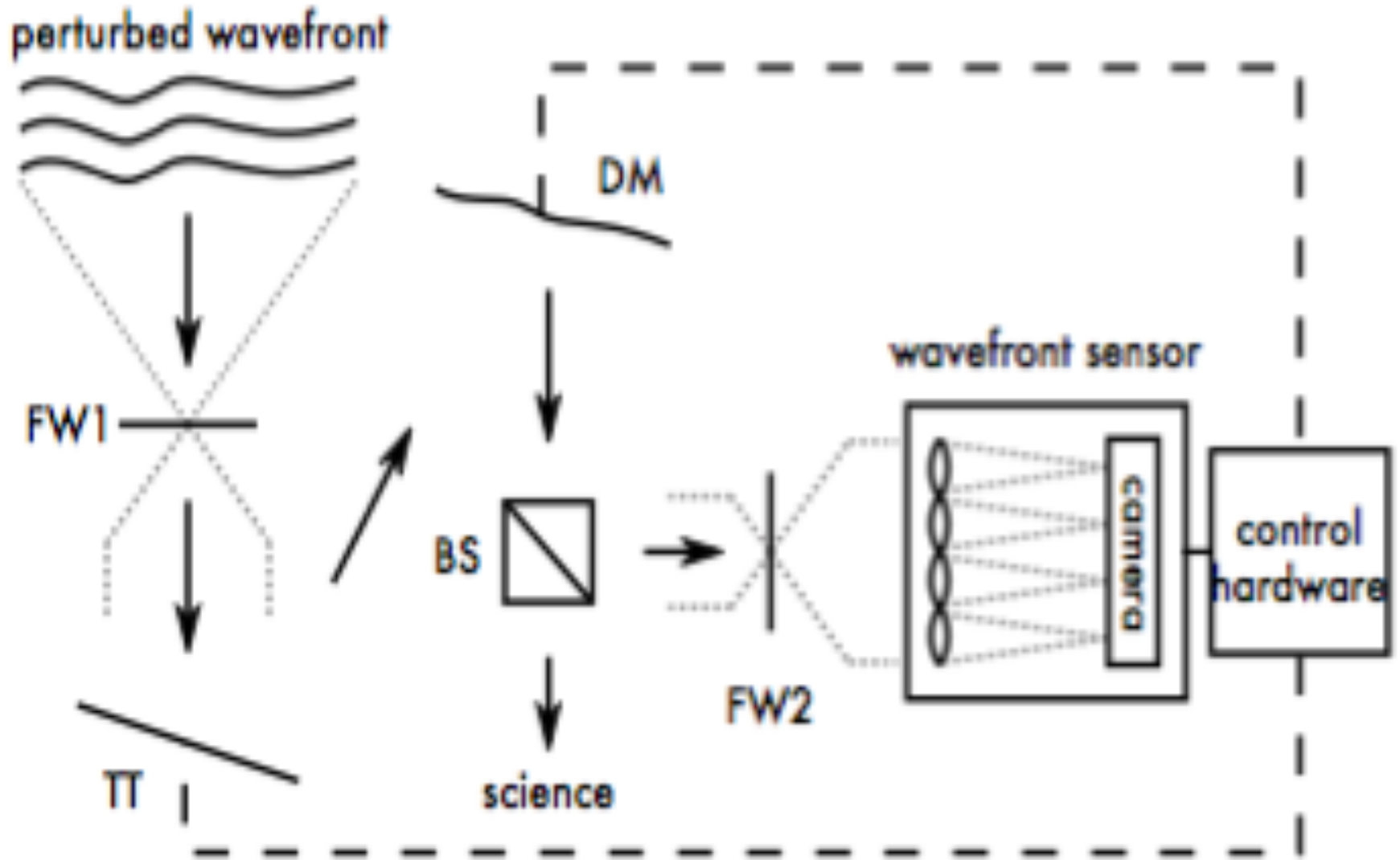
Lick
Observatory

Why Build a RAVEN Demonstrator ?

- AO – Successful on large telescopes, but small images (10 arcsecs). Good for Stars – Galaxies?
- GMT, E-ELT, TMT - need MOAO to be fully utilized.
- Raven demonstrated MOAO AND did science tasks.



AO – Key Elements

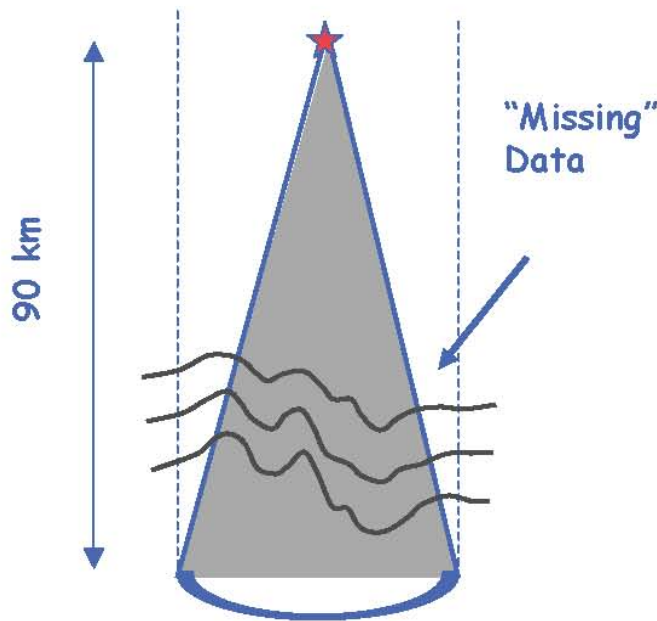


LGSコーン効果の低減: LTAO

AO tomography: measure turbulence and correct for “cone effect”

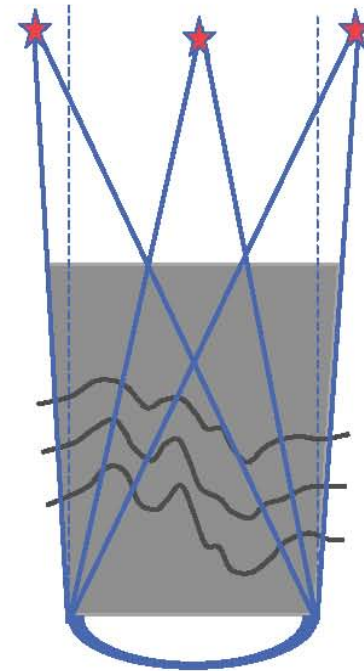


one laser guide star



Without tomography

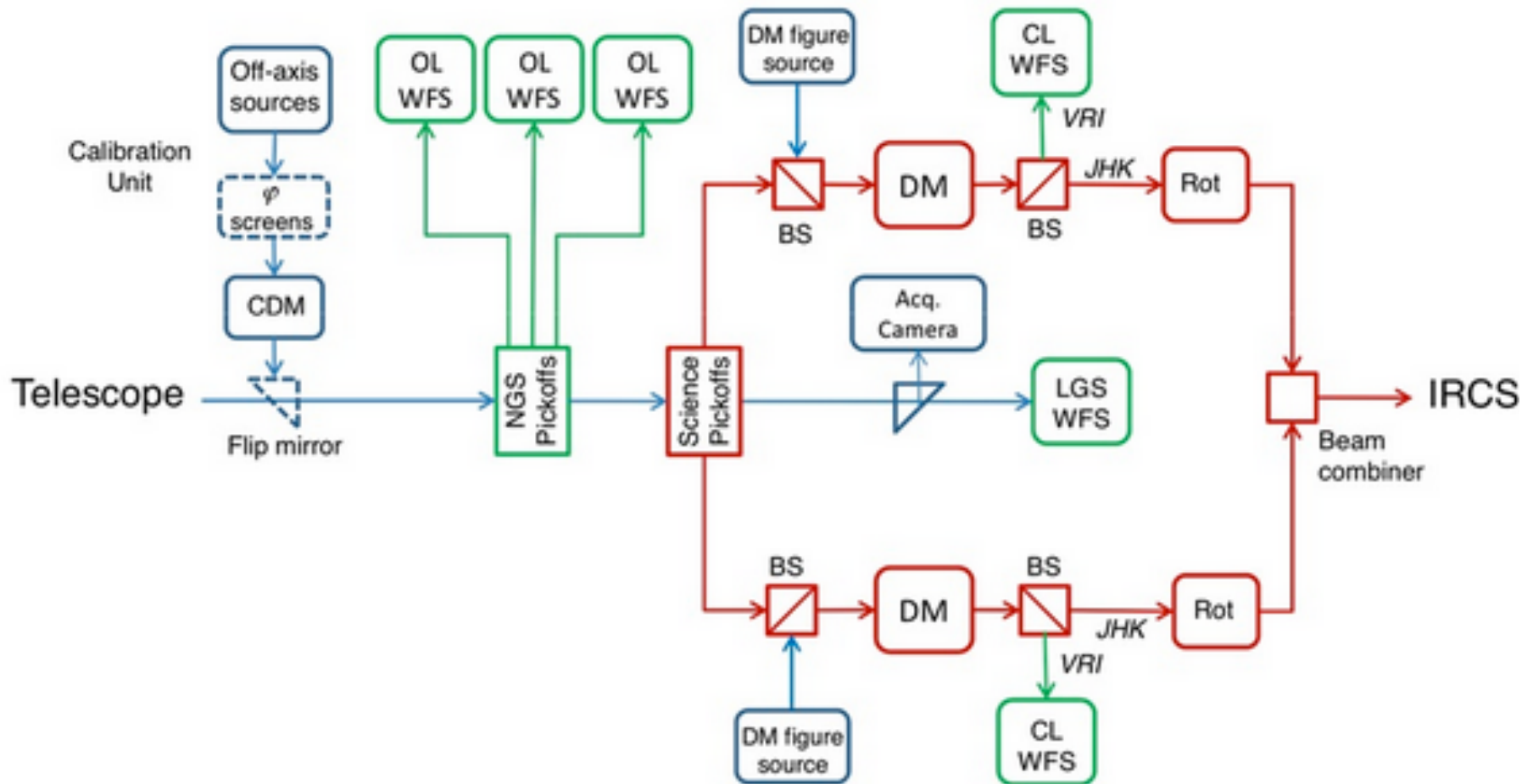
multiple laser guide stars



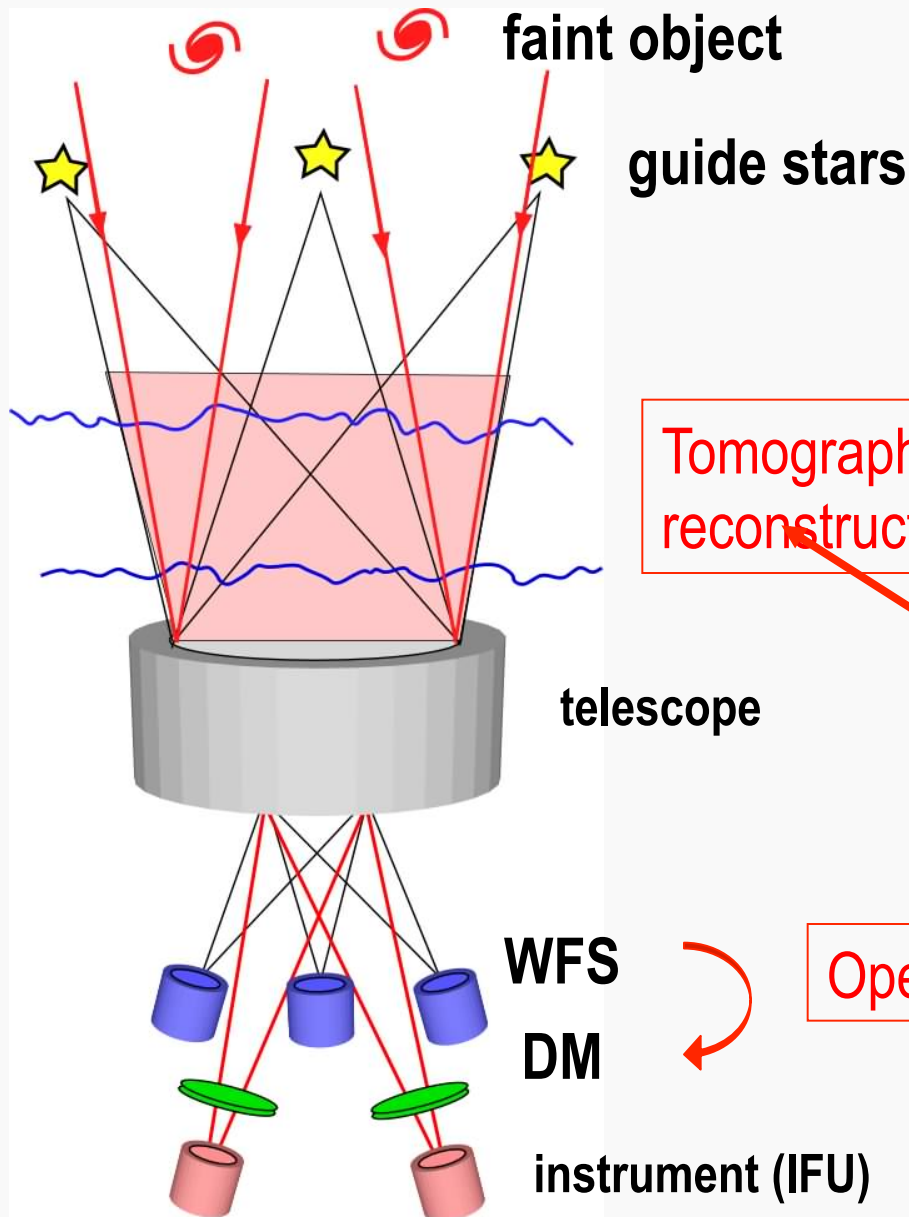
With tomography

With AO tomography, measure 3D distribution of turbulence above telescope

Raven System Block Diagram



MOAO: Multi-Object AO



FOV:

- patrol: \sim arcmin
- unit: \sim arcsec
- diffraction-limit
- targeted obs.

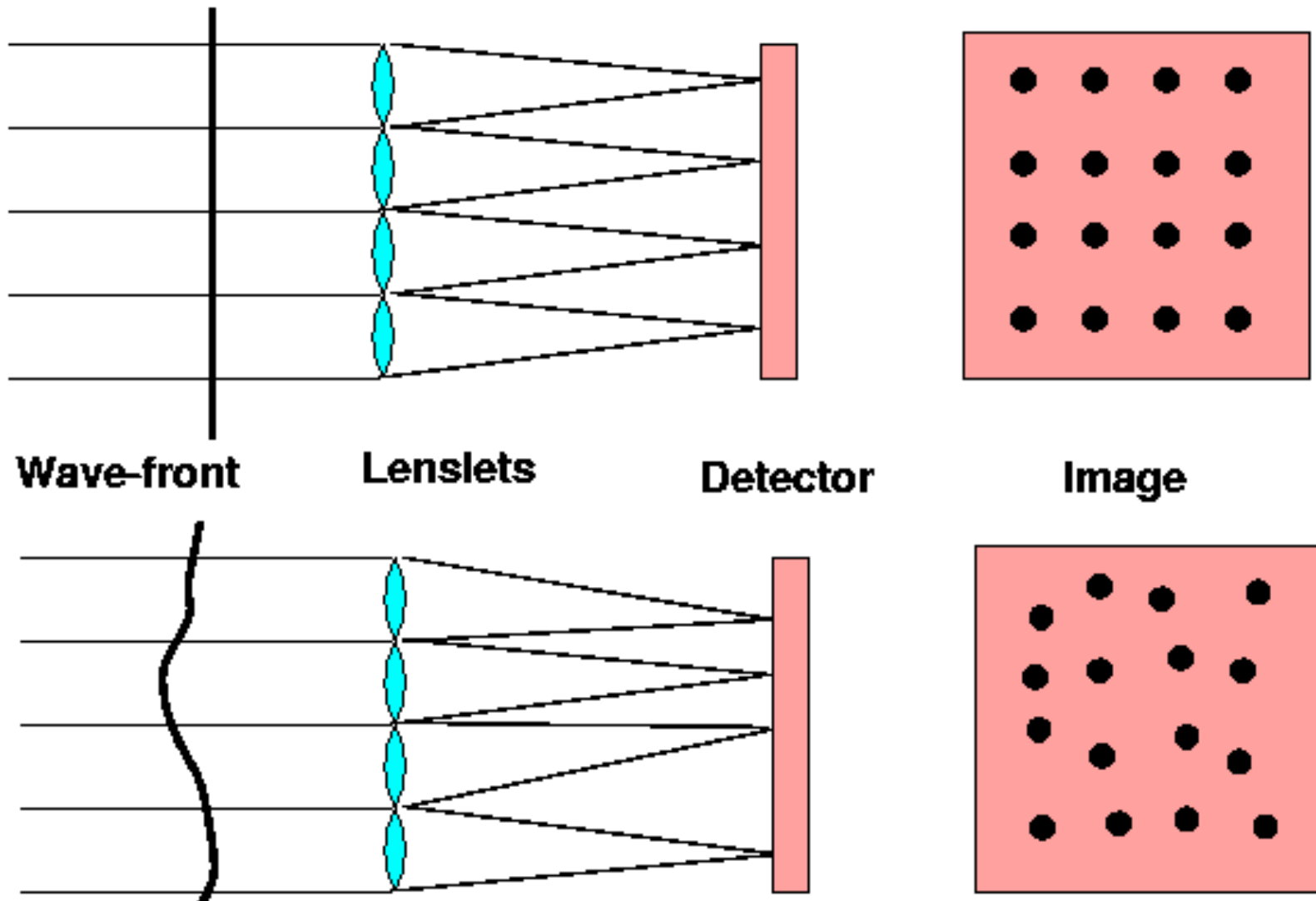
Tomographic wavefront reconstruction

Key Technology

Open-loop control

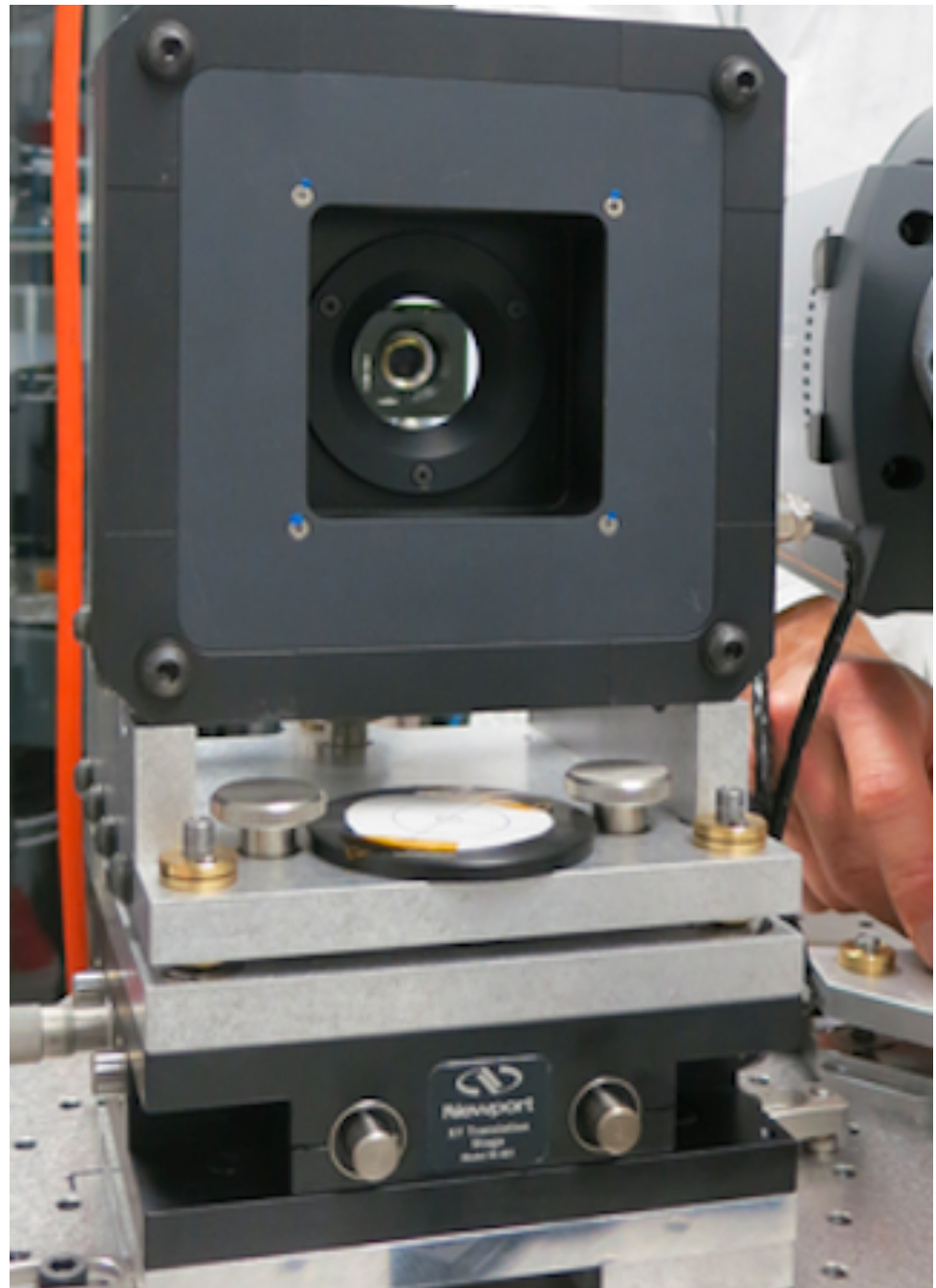
Calibration !!

WFS (Shack Hartman)

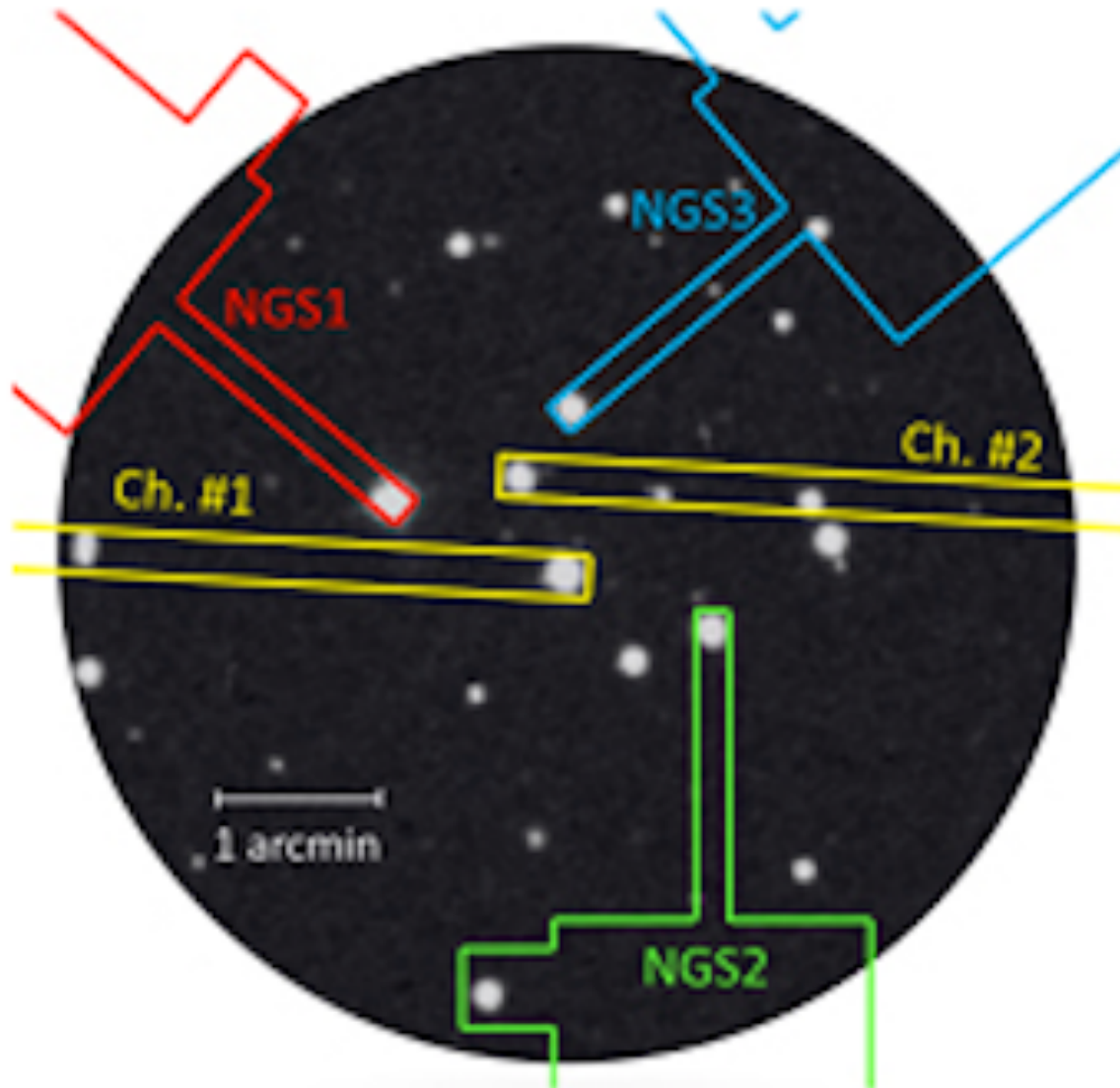


DM's

- Mirror Tilt/Tip (T/T)
- Mirror surface deformed by actuators (DM)
- Raven's DM's have both T/T and DM

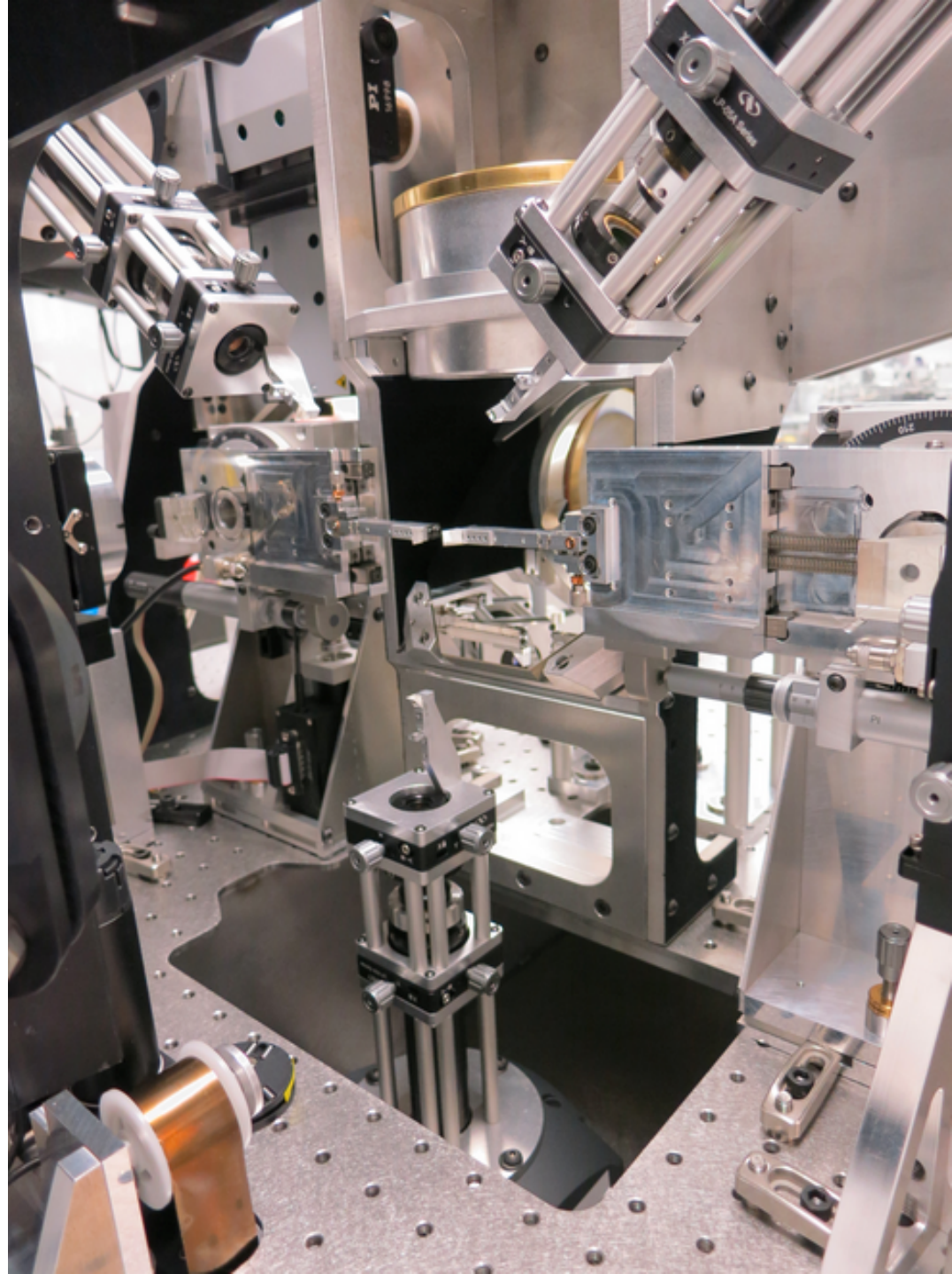


RAVEN Pick-Off arms in Light Beam

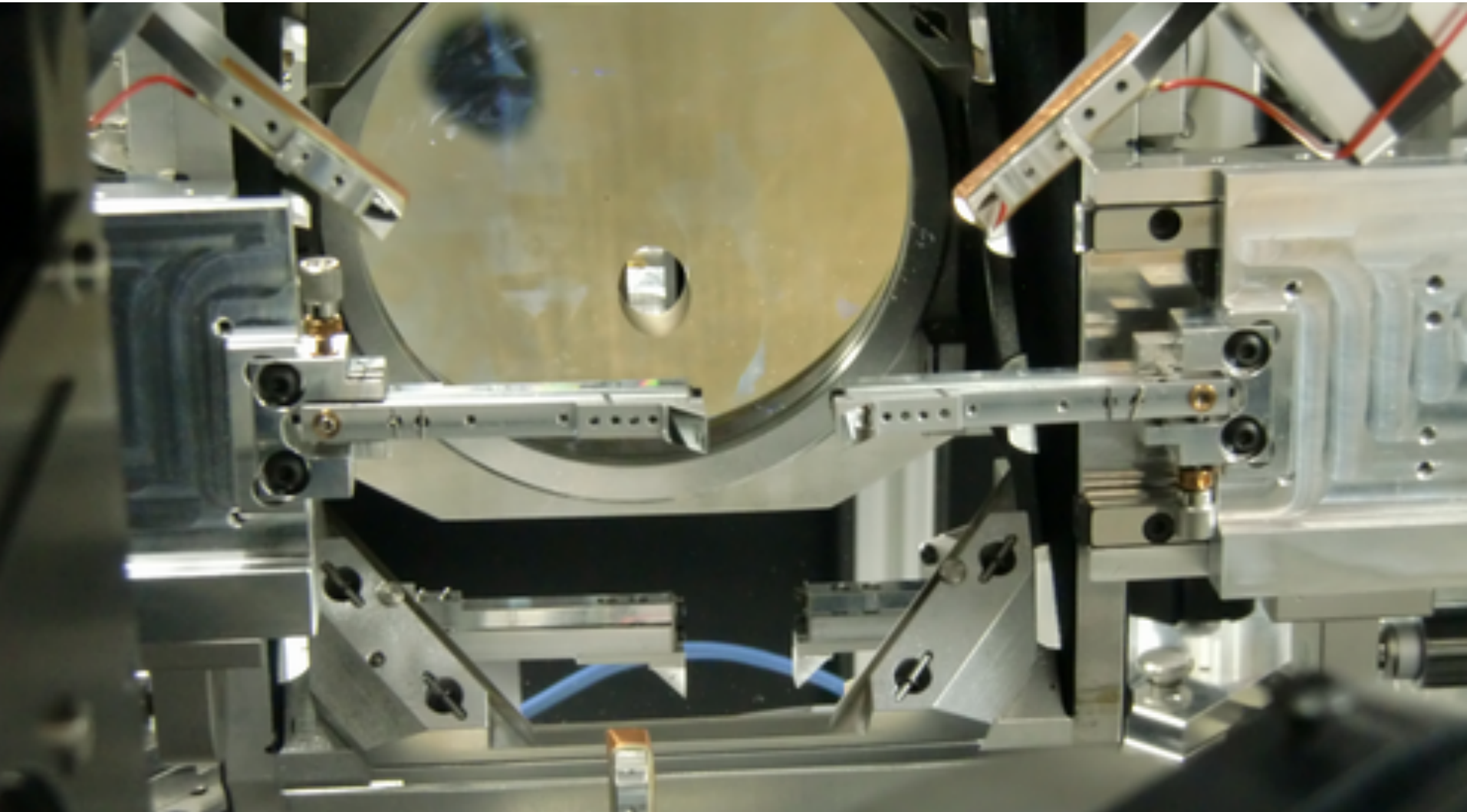


MOAO

- Multiple NGS's
- Multiple LGS's
- Multiple WFS's
- Multiple DM's
- Multiple Pick-off mirrors
- Multiple Science measurements

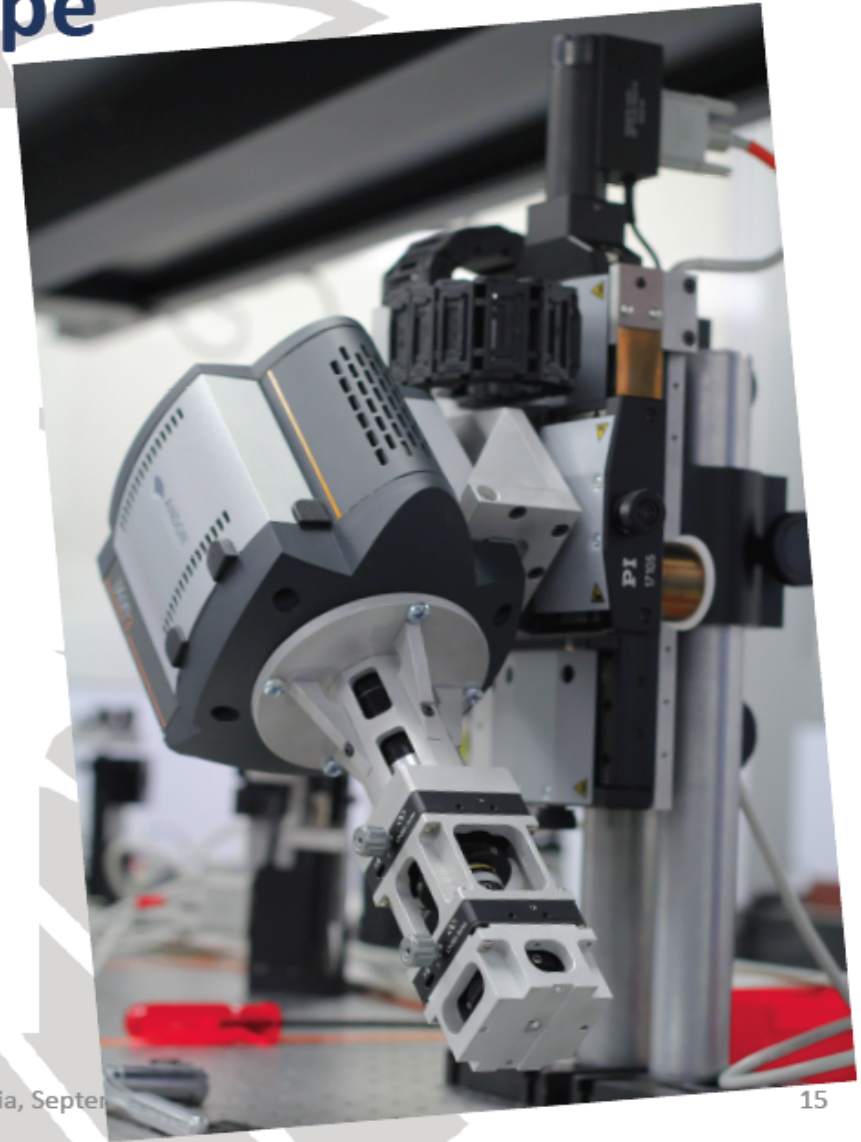
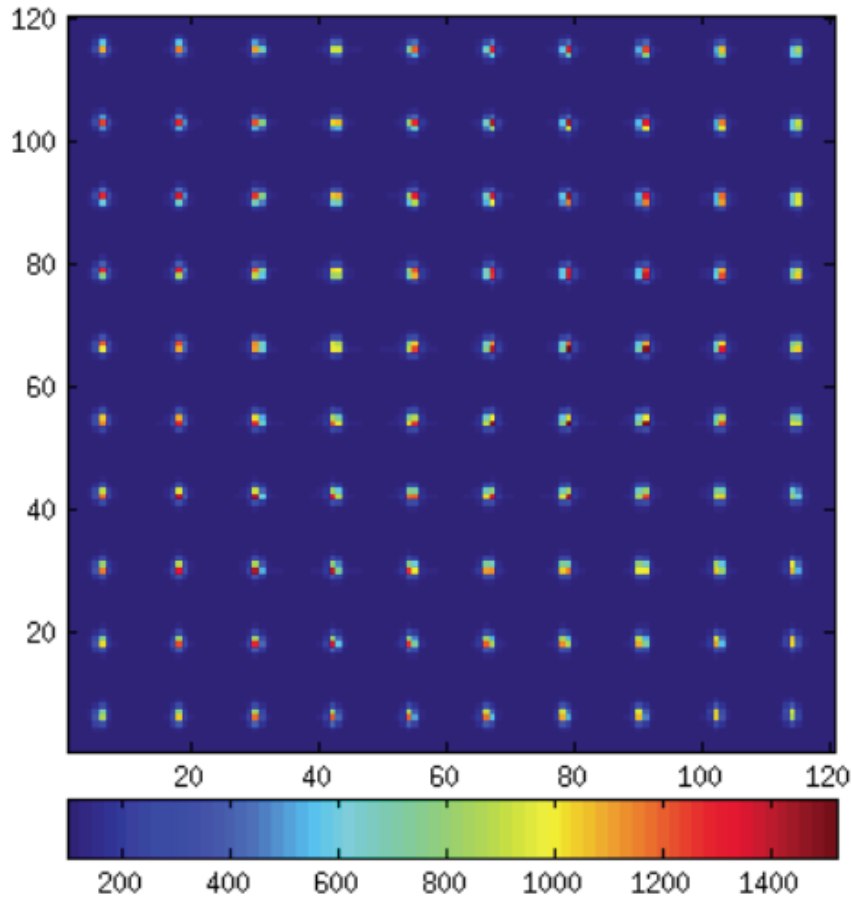


Pick Off Arms

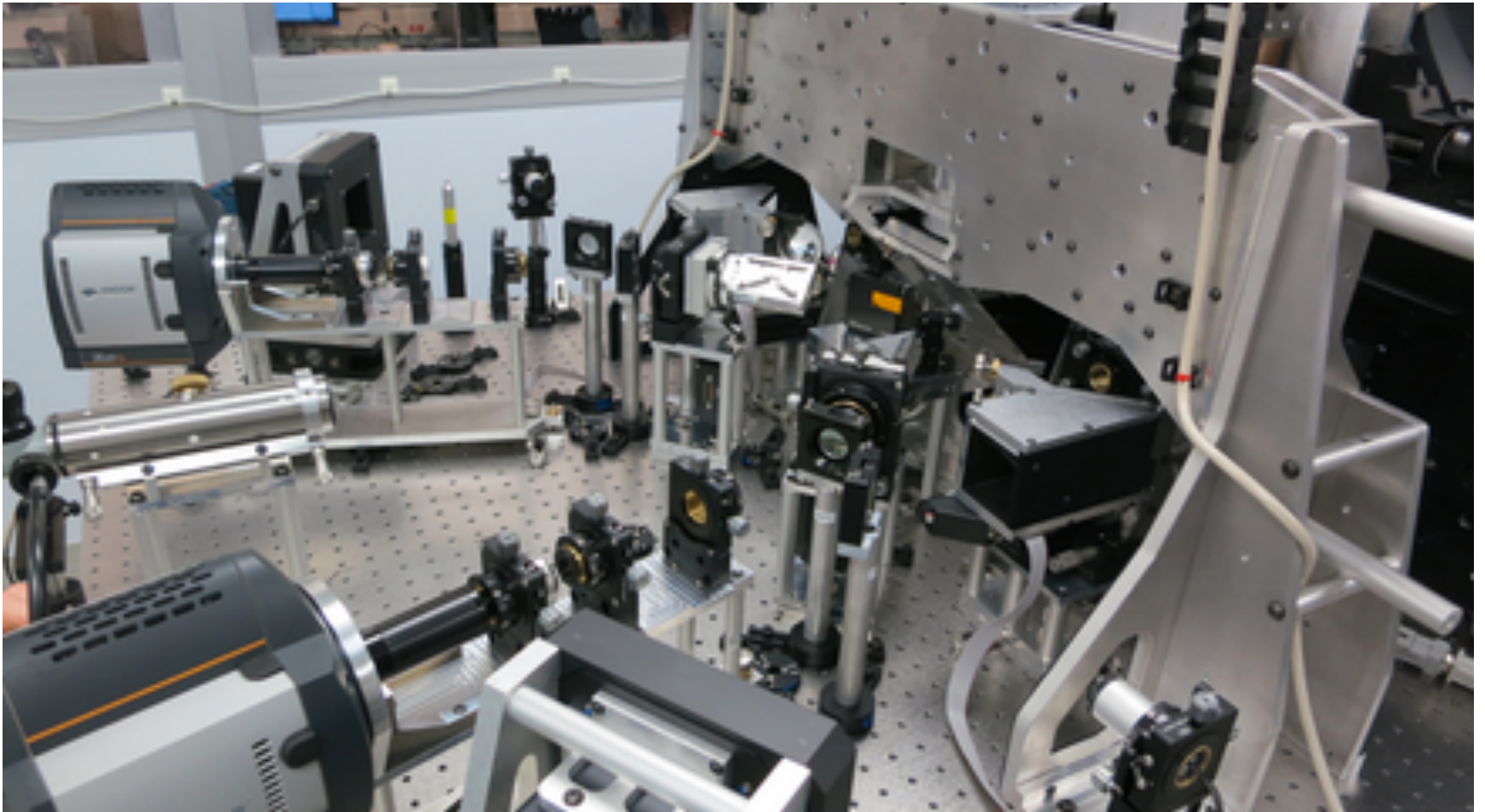




OL WFS Prototype



RAVEN optical bench

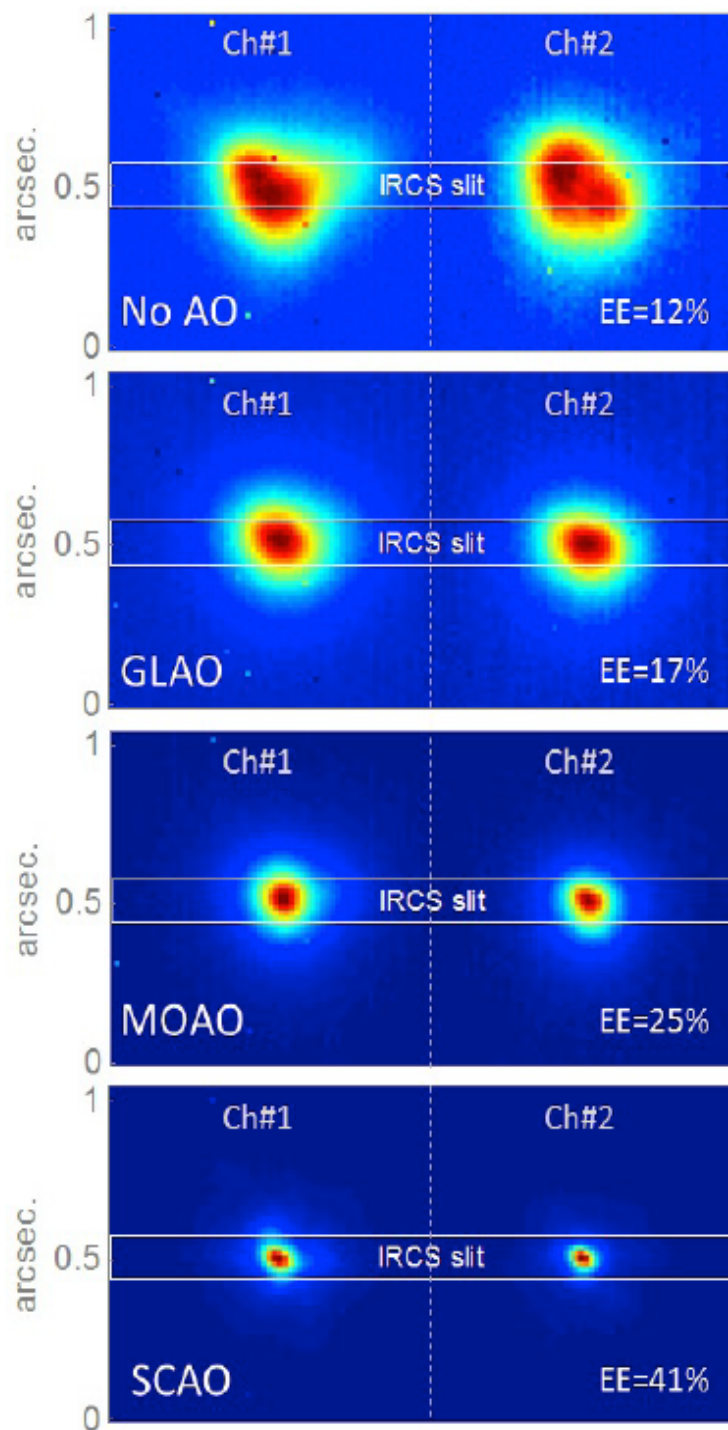


Final assembly





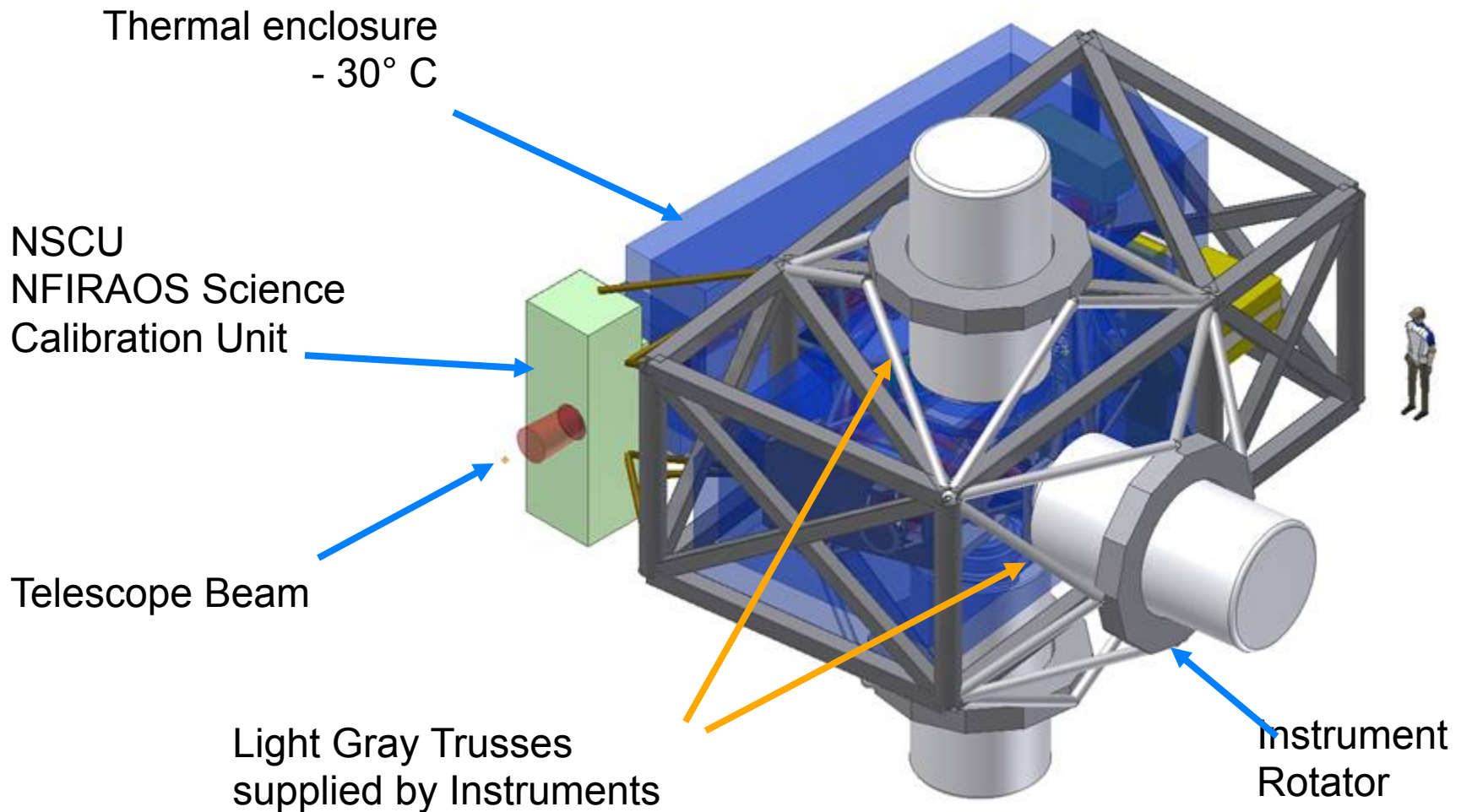
Results



Overview

- *One small target at a time limits observing efficiency.*
- *Next generation of large telescopes (GMT, E-ELT, & TMT) will need AO over large fields of view. Hence MOAO systems.*
- *Such MOAO systems “might have on the order of 8 LGS sensors and 20 science pick-off arms” (Andersen 2014).*
- *Raven has demonstrated feasibility.*
- *Raven did this and performed science tasks!*
- *HIA and UVIC are part of the TMT Team!*

Cooled Enclosure with Calibration Unit, 3 Instruments, and Support Structure



Questions? Comments!



The Raven Team

References

- Brent Ellerbroek, TMT Early Light Adaptive Optics, TMT @ SPIE 2010, San Diego, June 26, 2010, TMT.AOS.PRE.10.054.REL01
 - <http://web.uvic.ca/~ravenmoa/>
 - <http://web.uvic.ca/~ravenmoa/Blog.html>
 - <http://www.subarutelescope.org/Topics/2014/08/26/index.html>
 - <http://www.nrc-cnrc.gc.ca/eng/rd/nsi/>
 - <http://web.uvic.ca/lacir/optics/>
 - <http://planetimager.org/>
- <http://astroherzberg.org/projects/>
- *John Bochanski Next Generation AO*
<http://www.skyandtelescope.com/astronomy-news/next-gen-adaptive-optics-09092014/>
- <https://en.wikipedia.org/wiki/Twinkling>
-
- https://en.wikipedia.org/wiki/Deformable_mirror
-
- https://en.wikipedia.org/wiki/Microelectromechanical_systems
-
- https://en.wikipedia.org/wiki/Shack%E2%80%93Hartmann_wavefront_sensor
- https://en.wikipedia.org/wiki/Wavefront_sensor
-
- <http://www.ctio.noao.edu/~atokovin/tutorial/part3/wfs.html>