## Lecture 1 problems

- 1. The significance of a certain conclusion depends very strongly on whether the most luminous known quasar is included in the dataset. The object is legitimately in the dataset in terms of the pre-stated selection criteria. Is the conclusion robust?
- 2. We have N = 10 measurements of a variable  $x_i = (7.6, 5.8, 8.0, 6.9, 7.2, 7.5, 6.4, 8.1, 6.3, 7.0)$ . Estimate the mean, variance and median of this variable. What are the errors in your estimates?
- 3. I observe 100 galaxies, 30 of which are AGN. What is the best estimate of the AGN fraction and its error?
- 4. In the HST guide star catalogue, 60% of the objects are binary stars. How large a sample should be chosen to ensure that the probability of the sample containing at least 2 non-binary stars is at least 99%?
- 5. The density of quasars on the sky is known to be  $20 \text{ deg}^{-2}$ . What area of sky would we need to survey to ensure a 99% chance of finding a quasar?
- 6. A galaxy of absolute magnitude M = -20 is observed to have apparent magnitude  $m = 20.0 \pm 0.2$ . What is the luminosity distance  $D_L$  in Mpc, and its error? [Assume  $m M = 5 \log_{10} D_L + 25$ ].
- 7. The total mass of a binary star system (in solar masses) is given by Kepler's law  $M = a^3/P^2$ , where a is the mean separation in A.U. and P is the period in years. The  $\alpha$  Centauri system has a period of  $79.9 \pm 1.0$  years and mean separation  $a = 23.7 \pm 1.0$  AU. What is the total mass and its error?
- 8. We have N = 5 measurements of a quantity:  $(7.4 \pm 2.0, 6.5 \pm 1.1, 4.3 \pm 1.7, 5.5 \pm 0.8, 6.0 \pm 2.5)$ . What is the optimal estimate of this quantity and the error in that estimate? A further measurement  $3.0 \pm 0.2$  is added. How should our estimate change?