Science from overlapping lensing / spec-z surveys

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Probes of the cosmological model

How fast is the Universe expanding with time?

How fast are structures growing within it?





Redshift-space distortions

 RSD allow spectroscopic galaxy surveys to measure the growth rate of structure coherent



Why combination of lensing and RSD?

- Sensitive to theories of gravity in complementary ways
- General perturbations to FRW metric:

$$ds^2 = \left[1 + 2\psi(x,t)\right] dt^2 - a^2(t) \left[1 - 2\phi(x,t)\right] dx^2$$

- (ψ, ϕ) are metric gravitational potentials, identical in General Relativity but can differ in general theories
- Relativistic particles (e.g. light rays for lensing) collect equal contributions and are sensitive to $(\psi+\phi)$
- Non-relativistic particles (e.g. galaxies infalling into clusters) experience the Newtonian potential ψ

Applications



Overlaps of lensing and spec-z surveys

- Improvement of cosmological measurements through addition of galaxy-galaxy lensing
- [e.g. determines bias of lens sample which improves RSD measurements of lenses, especially when using multiple-tracer techniques, e.g. Cai & Bernstein (2012)]
- Spec-z survey allows definition of lens samples (e.g. groups, galaxy types) enabling a range of studies
- Understanding, calibration and risk mitigation of systematic errors (photo-z errors including outliers, intrinsic alignments, cosmic shear)

Overlaps of lensing and spec-z surveys

• Mis-match between imaging and spectroscopic surveys!



BOSS DR10 SGP



Overlaps of lensing and spec-z surveys

 Many recent papers considering impact for cosmology of same-sky vs. different-sky lensing/spec-z surveys



Photometric redshift calibration

- Photometric redshift errors are one of the leading systematics for weak lensing tomography
- Mean and width of redshift distributions in each photo-z bin must be known to accuracy ~ 10⁻³
- Method (I) : spectroscopic training set [issues : sample variance, incompleteness of training set, outliers]
- Method (2) : photo-z/spec-z cross-correlations [issues : degeneracies with galaxy bias, cosmic magnification]
- Can OzDES currently help with either method?

Photometric redshift calibration

- Training set method : suppressing sample variance systematic for DES weak lensing requires 100 AAOmega pointings or 5 VIPERS surveys !
- Phot-z/spec-z cross-correlation method : need ~50,000 spec-z's per unit redshift -- can use 2dFGRS, SDSS, BOSS at z < 0.6 and VIPERS at z > 0.6?
- OzDES is not currently transformational for this science, but photo-zs are still useful for other large-scale structure topics with less stringent requirements

Future AAT proposals?

- OzDES-wide proposed in March 2013 (150 nights, 3000 deg² coverage) not approved
- Issue (1) : time now allocated to other projects (SAMI, OzDES-deep) - only 10-15 nights/semester remain?
- Issue (2) : can only observe DES fields in B semesters
- Issue (3) : need to map > 1000 deg² to be competitive (e.g. existing RCS2, future HSC projects)
- Current status : musing on competitiveness of 2014B proposal targetting KiDS, later widening to DES?